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**Magma Plumbing Systems: A Geophysical Perspective**

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**Abstract**

Over the last few decades, significant advances in using geophysical techniques to image the structure of magma plumbing systems have enabled the identification of zones of melt accumulation, crystal mush development, and magma migration. Combining advanced geophysical observations with petrological and geochemical data has arguably revolutionised our understanding of and afforded exciting new insights into the development of entire magma plumbing systems. However, divisions between the scales and physical settings over which these geophysical, petrological, and geochemical methods are applied still remain. To characterise some of these differences and promote the benefits of further integration between these methodologies, we provide a review of geophysical techniques and discuss how they can be utilised to provide a structural context for and place physical limits on the chemical evolution of magma plumbing systems. For example, we examine how Interferometric Synthetic Aperture Radar (InSAR), coupled with Global Positioning System (GPS) and Global Navigation Satellite System (GNSS) data, and seismicity may be used to track magma migration in near real-time. We also discuss how seismic imaging, gravimetry, and electromagnetic data can image contemporary melt zones, magma reservoirs, and/or crystal mushes. These techniques complement seismic reflection data and rock magnetic analyses that delimit the structure and emplacement of ancient magma plumbing systems. For each of these techniques, with the addition of full-waveform inversion (FWI), the use of Unmanned Aerial Vehicles (UAVs), and the integration of geophysics with numerical modelling, we discuss potential future directions. We show that approaching problems concerning magma plumbing systems from an integrated petrological, geochemical, and geophysical perspective will undoubtedly yield important scientific advances, providing exciting future opportunities for the volcanological community.

## **1. Introduction**

Igneous petrology and geochemistry are concerned with the chemical and physical mechanisms governing melt genesis, mobilisation, and segregation, as well as the transport/ascent, storage,

54 evolution, and eruption of magma. The reasons for studying these fundamental processes include  
55 understanding volcanic eruptions, modelling the mechanical development of magma conduits and  
56 reservoirs, finding magma-related economic ore deposits, exploring for active geothermal energy  
57 sources, and determining the impact of magmatism in different plate tectonic settings on the  
58 evolution of the lithosphere and crustal growth. However, whilst petrological and geochemical  
59 studies over the last century have shaped our understanding of the physical and chemical evolution  
60 of magma plumbing systems, assessing the distribution, movement, and accumulation of magma in  
61 the Earth's crust from these data remains challenging. A key frontier in igneous petrological and  
62 geochemical research thus involves deciphering how and where magma forms, the routes it takes  
63 toward the Earth's surface, and where exactly it is stored.

64         This contribution will demonstrate how geophysical data can be used to determine the  
65 architecture of magma plumbing systems, providing a structural framework for the interpretation of  
66 petrological and geochemical data. To aid the alignment of petrological, geochemical, and  
67 geophysical disciplines it is first important to delineate what we mean by 'magma'. We follow  
68 Glazner *et al.*, (2016) and define magma as, "naturally occurring, fully or partially molten rock  
69 material generated within a planetary body, consisting of melt with or without crystals and gas  
70 bubbles and containing a high enough proportion of melt to be capable of intrusion and extrusion".  
71 Importantly, this definition specifically considers that magma: (i) forms through the migration and  
72 accumulation of partial melt that is initially distributed throughout pore spaces in a rock volume;  
73 and (ii) is a suspension of particles (i.e. crystals, xenoliths, and/or bubbles) within melt (see  
74 Cashman *et al.*, 2017). As magma starts to solidify, the proportion of suspended crystals and thus  
75 the relative viscosity of the magma increases until a relatively immobile, continuous network of  
76 crystals and interstitial melt develops; we term this a 'crystal mush' (e.g., Hildreth, 2004; Glazner *et*  
77 *al.*, 2016; Cashman *et al.*, 2017). The rheological transition from a magma to a crystal mush is  
78 partly dependent on its chemistry, but typically occurs abruptly when the particle volume increases  
79 across the 50–65% range (Cashman *et al.*, 2017). Crystal mushes thus exist at or above the solidus

80 and generally cannot be erupted, although they may be partly entrained in eruptible magma as  
81 glomerocrysts, cumulate nodules, or restite (Cashman *et al.*, 2017). Migration of interstitial melt  
82 within a crystal mush can lead to its accumulation and, thus, formation of a magma. A magma  
83 plumbing system therefore consists of interconnected magma conduits and reservoirs, which store  
84 magma as it evolves into a crystal mush, ultimately fed from a zone of partial melting (e.g., Fig. 1).  
85 These definitions are supported by geophysical imaging and analyses of contemporary reservoirs,  
86 which show melt volumes in the mid- to upper crust are typically low (<10%) and likely exist  
87 within a crystal mush (e.g., Paulatto *et al.*, 2010; Koulakov *et al.*, 2013; Ward *et al.*, 2013;  
88 Hammond, 2014; Comeau *et al.*, 2015; Comeau *et al.*, 2016; Delph *et al.*, 2017). These definitions  
89 and geophysical data question the traditional view that magma resides in long-lived, liquid-rich, and  
90 volumetrically significant magma chambers. Following this, the emerging paradigm for igneous  
91 systems is thus that liquid-rich magma chambers are short-lived, transient phenomena with: (i) melt  
92 typically residing in mushes that develop through the incremental injection of small, distinct magma  
93 batches; and (ii) magma accumulating in thin lenses (e.g., Hildreth, 2004; Annen *et al.*, 2006;  
94 Annen, 2011; Miller *et al.*, 2011; Solano *et al.*, 2012; Cashman & Sparks, 2013; Annen *et al.*, 2015;  
95 Cashman *et al.*, 2017). We are now starting to view magmatic systems as a vertically extensive,  
96 transcrustal, interconnected networks of magma conduits and magma/mush reservoirs (Fig. 1) (e.g.,  
97 Cashman *et al.*, 2017).

98         The current use of geophysical techniques within the igneous community can be separated  
99 into two distinct areas focused on either characterising active volcanic domains or investigating the  
100 structure and emplacement of ancient magma plumbing systems. For example, in areas of active  
101 volcanism, our understanding of magma plumbing system structure principally comes from the  
102 application of geophysical techniques that detect sites of magma movement or accumulation (e.g.,  
103 Sparks *et al.*, 2012; Cashman & Sparks, 2013). Such geophysical techniques include Interferometric  
104 Synthetic Aperture Radar (InSAR; e.g., Biggs *et al.*, 2014), seismicity (e.g., recording of  
105 earthquakes associated with magma movement; e.g., White & McCausland, 2016), various seismic

106 imaging methods (e.g., Paulatto *et al.*, 2010; Hammond, 2014), gravimetry (e.g., Battaglia *et al.*,  
107 1999; Rymer *et al.*, 2005), and electromagnetic techniques (Desissa *et al.*, 2013; Comeau *et al.*,  
108 2015). These techniques allow examination of: (i) the temporal development of magma plumbing  
109 systems (e.g., Pritchard & Simons, 2004; Sigmundsson *et al.*, 2010); (ii) vertical and lateral  
110 movements of magma (e.g., Keir *et al.*, 2009; Jay *et al.*, 2014); (iii) the relationship between  
111 eruption dynamics, volcano deformation, and intrusion (e.g., Sigmundsson *et al.*, 2010;  
112 Sigmundsson *et al.*, 2015); and (iv) estimates of melt sources and melt fractions (e.g., Desissa *et al.*,  
113 2013; Johnson *et al.*, 2016). However, inversion of these geophysical data typically results in non-  
114 unique, relatively low-resolution models of subsurface structures. Furthermore, some methods only  
115 capture active processes, which may be short-lived or even instantaneous, potentially providing  
116 information on only a small fraction of the magma plumbing system.

117 In contrast to the study of active volcanic domains, the analysis of ancient plumbing systems  
118 through field observations, geophysical imaging techniques (e.g., reflection seismology, gravity,  
119 and magnetic data), and/or rock magnetic experiments can provide critical insights into magma  
120 emplacement, mush evolution, and allow the geometry of entire plumbing systems to be  
121 reconstructed (e.g., Cartwright & Hansen, 2006; Stevenson *et al.*, 2007a; Petronis *et al.*, 2013;  
122 Muirhead *et al.*, 2014; O'Driscoll *et al.*, 2015; Magee *et al.*, 2016). Whilst such studies of ancient  
123 plumbing systems provide a framework for interpreting the structure of active intrusion networks,  
124 capturing a snapshot of how magma moved and melt was distributed through the system at any one  
125 time is difficult because magmatism has long since ceased.

126 All the techniques employed to define active and ancient plumbing systems, including  
127 petrological and chemical analyses, provide information at different spatial and/or temporal  
128 resolutions. Answering the major outstanding questions in studies of magma plumbing systems  
129 therefore requires the integration of complementary petrological, geochemical, geophysical,  
130 geochronological, and structural techniques. Here, we examine active plumbing systems using  
131 InSAR, seismicity, seismic imaging, gravimetry, and electromagnetic techniques. To provide a

context for the interpretation of data pertaining to the active systems, we also discuss how seismic reflection data and rock magnetic techniques can be used to derive the structure and evolution of ancient intrusion networks. The potential of emerging techniques involving seismic full-waveform inversion (FWI) and unmanned aerial vehicles (UAVs) are also considered, as is the role of numerical modelling in bringing together outputs from different datasets. For each technique described, we briefly discuss the methodology and limitations and provide a summary of the key findings and potential uses, with a focus on integration with petrological and geochemical data. The aim of this review is to facilitate and promote integration between petrologists, geochemists, geochronologists, structural geologists, and geophysicists interested in addressing outstanding problems in studies of magma plumbing systems.

142

## 143 **2. Understanding magma plumbing system structure**

Here, we discuss a range of techniques that can be utilised to define different aspects of magma plumbing system structure and evolution. In particular, we describe how InSAR, seismicity, seismic imaging (e.g., seismic tomography), gravity, and electromagnetic data is used to determine melt fractions and distribution, track movement of magma in near real-time, and/or locate sites and examine the evolution of magma/mush storage. Overall, these geophysical techniques allow the structure of active plumbing systems and their transient evolution to be assessed. We also discuss how seismic reflection data can provide unprecedented images of ancient plumbing systems and associated host rock deformation in three-dimensions at resolutions of 10's of metres. Finally, we examine the application of rock magnetic techniques to assess magma flow and crystallisation processes at a range of scales.

Although beyond the scope of this review, it is critical to highlight that interpreting the geophysical response of a rock or magma relies on understanding its physical and chemical properties (e.g., density, temperature, and melt fraction). Laboratory experiments testing how rock or magma properties influence geophysically measured parameters (e.g., seismic velocities and

158 resistivity) thus provide context for interpreting magma plumbing system structure and evolution  
159 from geophysical data (e.g., Gaillard, 2004; Pommier *et al.*, 2010; Pommier, 2014).

160

## 161 **2.1. Insights into magma plumbing systems from ground deformation data**

### 162 *Technique*

163 Changes in volume within magma plumbing systems can deform the host rock, potentially resulting  
164 in displacement of the Earth's surface. Such displacements are a unique source of information for  
165 volcanologists and can be modelled to estimate geodetic source depth and, to varying extents, the  
166 source geometry and volume change (e.g., Segall, 2010). Measuring the deformation of the Earth's  
167 surface can thus provide information about the characteristics and timing of magma movement and  
168 accumulation, as well as variations in internal reservoir conditions. Traditionally, deformation  
169 measurements are made using levelling, electronic distance meters, tiltmeters, and Global  
170 Positioning System (GPS), all of which have proven to be reliable methods and thus are widely  
171 used in volcano monitoring (e.g., Dzurisin, 2006). For example, GPS measurements retrieve the  
172 relative positions of receivers on Earth's surface from dual frequency carrier phase signals  
173 transmitted from GPS or Global Navigation Satellite System (GNSS) satellites with precisely  
174 known orbits. Distances between satellites and receivers are assessed from the travel-time, i.e. the  
175 measured difference between the transmitted and received times of a unique ranging code, allowing  
176 movement of the Earth's surface over time to be monitored (see review by Dixon, 1991).  
177 Permanently installed receivers record position data continuously, but receivers can also be  
178 deployed for a limited time during GPS campaigns to provide additional measurements, normally  
179 made relative to a standard benchmark location (e.g., Dvorak & Dzurisin, 1997). Whilst tiltmeters  
180 and GPS can provide continuous measurements, their spatial resolution is limited by logistical  
181 constraints such as cost and accessibility, which may be restricted at active volcanoes.

182 The geographic reach of volcano geodesy has been greatly expanded over the past two  
183 decades by the application of Interferometric Synthetic Aperture Radar (InSAR), an active remote



184 sensing technique that uses microwave electromagnetic radiation to image the Earth's surface (e.g.,  
185 Simons & Rosen, 2007; Pinel *et al.*, 2014). Surface displacements can be measured by constructing  
186 interferograms, where the difference in phase between radar echoes from time-separated images  
187 appear as 'fringes' of variation in the line of sight distance to the satellite (Fig. 2). The patterns of  
188 fringes in individual interferograms are distinctive for different deformation source geometries,  
189 such as for horizontal (sill-like) or vertical (dyke-like) opening of intrusions, or the pressurisation of  
190 a spheroidal reservoir (i.e. a Mogi source) (e.g., Fig. 2b). However, magma intrusion processes can  
191 rarely be uniquely identified from geodetic source geometry alone, and distinguishing between  
192 magmatic, hydrothermal, structural (e.g., faulting and compaction), and combinations of elastic and  
193 inelastic sources is particularly challenging (e.g., Galland, 2012; Holohan *et al.*, 2017).

194         Whilst a single interferogram only provides displacements in satellite line-of-sight, a  
195 pseudo-3D displacement field can be estimated by combining multiple images from polar orbits that  
196 are ascending (i.e. satellite moves roughly northward, looking east) and descending (i.e. satellite  
197 moves roughly southward, looking west) (Fig. 2a), especially where GNSS measurements can also  
198 be incorporated. The lateral spatial resolution of most InSAR data is on the order of metres to tens  
199 of metres, whilst vertical movements can be resolved on the order of centimetres and sometimes  
200 millimetres. Temporal resolution depends on the satellite revisit time and ranges between days to  
201 months depending upon the sensor type and satellite orbit. This means that InSAR can be used to  
202 regularly assess ground deformation at virtually any volcano worldwide situated above sea level,  
203 with a higher spatial density of measurements than achieved using from ground-based  
204 instrumentation.

205         Magmatic processes are only observable by InSAR when either magma movement or  
206 internal reservoir processes (e.g., cooling and contraction, phase changes) cause changes in pressure  
207 and thereby instigate deformation of the host rock and free surface. The best-fit parameters of a  
208 deformation source (e.g., an intruding magma body) are most often assessed by inverting measured  
209 displacements using analytical elastic-half space models of simple source geometries, although

210 there are often trade-offs between parameters such as source depth and volume change (e.g.,  
211 Pritchard & Simons, 2004). Complex and more realistic deformation source geometries may be  
212 retrieved using finite element-based linear inversion of displacement fields (e.g., Ronchin *et al.*,  
213 2017). A proportion of any pressure change may be accommodated by magma compressibility,  
214 leading to underestimation of volume changes (e.g., Rivalta & Segall, 2008; McCormick-Kilbride *et*  
215 *al.*, 2016). Assessing both volume changes and especially the total volume of a magma reservoir  
216 from geodetic data therefore remains challenging. Furthermore, host rocks in areas of repeated  
217 intrusion that have been heated above the brittle-ductile transition are better described by a  
218 viscoelastic rheology (e.g., Newman *et al.*, 2006; Yamasaki *et al.*, 2018), while ductile  
219 accommodation of volume changes may occur at greater depth. Where some constraints are  
220 available for the structure and rheology of Earth's crust, finite or boundary element models may  
221 achieve a more realistic model of the deformation source (e.g., Masterlark, 2007; Hickey *et al.*,  
222 2017; Gottsmann *et al.*, 2017).

223

## 224 ***Observations***

225 Measurements of volcano deformation preceding and/or accompanying eruption have provided  
226 insights into the extent and structure of magma plumbing systems and, in some instances, the  
227 dynamics of magma movement. For example, InSAR-based observations at Eyjafjallajökull,  
228 Iceland have recognised the intrusion of multiple, distinct sills over a decade and their subsequent  
229 extraction when tapped during an explosive eruption (e.g., Pedersen & Sigmundsson, 2006;  
230 Sigmundsson *et al.*, 2010). Extensive lateral connections via dykes and sills between reservoirs  
231 and/or volcanoes have been illuminated by eruptions or unrest accompanied by ground deformation  
232 tens of kilometres away, and by the existence of multiple deformation sources (e.g., Alu-Dalafilla  
233 shown in Figures 3 and b, Pagli *et al.*, 2012; Korovin, Lu & Dzurisin, 2014; Cordon-Caulle, Jay *et*  
234 *al.*, 2014; Kenyan volcanoes, Biggs *et al.*, 2014; global synthesis, Ebmeier *et al.*, 2018). Inter-  
235 eruptive deformation at calderas is especially complex and seems to be particularly frequent and

236 high magnitude (e.g., Laguna del Maule; Fournier *et al.*, 2010; Singer *et al.*, 2014; Le Mével *et al.*,  
237 2015), with the location of the deformation sources inferred to vary over time (e.g., Campi Flegrei,  
238 Trasatti *et al.*, 2004; Yellowstone, Wicks *et al.*, 2006). The geometries of dykes and sills inferred  
239 from InSAR data inform our understanding of changing subsurface stress fields (e.g., Afar,  
240 Hamling *et al.*, 2010; Fernandina, Bagnardi *et al.*, 2013), as do measurements of displacements  
241 caused by moderate earthquakes in close proximity to magma plumbing systems (e.g., Kilauea,  
242 Wauthier *et al.*, 2013; Chiles-Cerro Negro, Ebmeier *et al.*, 2016).

243 At a transcrustal scale, deformation measurements have contributed to evidence for temporal  
244 variations in magma supply rates (e.g., in Hawaii, Poland *et al.*, 2012). Volume increases in the  
245 mid- to lower-crust, notably in the Central Andes, have provided the first observations of deep  
246 pluton growth (Pritchard & Simons, 2004). Furthermore, uplift during episodes of unrest that have  
247 not (yet) resulted in eruption have been detected at a broad range of volcanoes (e.g., Westdahl,  
248 Mount Peulik, Lu & Dzurisin, 2014; Alutu and Corbetti, Biggs *et al.*, 2011) and, in some cases,  
249 have been interpreted as evidence for the ‘pulsed’ accumulation of potentially eruptible magma  
250 (e.g., Santorini, Parks *et al.*, 2012). In addition to magma movement, volume changes associated  
251 with internal reservoir processes can also cause deformation of the host rock and free surface. For  
252 example, InSAR measurements have recorded subsidence linked to cooling and crystallisation of  
253 sills (Medicine Lake, Parker, 2016; Taupo Volcanic Zone, Hamling *et al.*, 2015). Transient periods  
254 of subsidence during inter-eruptive uplift have been attributed to phase transitions in response to the  
255 addition of more juvenile magma (e.g., Okmok, Caricchi *et al.*, 2014).

256

## 257 ***Implications and integration***

258 InSAR has increased the number of volcanoes where measurements of ground deformation have  
259 been made, from less than 50 in the late 1990s to over 200 today (Biggs & Pritchard, 2017; Ebmeier  
260 *et al.*, 2018). This increase in coverage has been particularly influential in the developing world  
261 where monitoring infrastructure is typically poor (Ebmeier *et al.*, 2013; Chaussard *et al.*, 2013),

262 with InSAR often providing the first evidence of magmatic activity at many volcanoes previously  
263 considered to be inactive (e.g., Pritchard & Simons, 2004; Biggs *et al.*, 2009; Biggs *et al.*, 2011; Lu  
264 & Dzurisin, 2014). A continued increase in the number and range of satellite- and large-scale UAV-  
265 based SAR instruments, as well as enhancements to their spatial and temporal resolution, will allow  
266 the detection of a greater range of volcanic ground deformation (e.g., Salzer *et al.*, 2014; Schaefer *et*  
267 *al.*, 2015; Stephens *et al.*, 2017). Overall, improved InSAR coverage will also increase the number  
268 of volcanoes where deformation measurements have been made across multiple cycles of eruption  
269 and deformation, increasing its usefulness for both hazard assessment and for characterising the  
270 extent, geometry, and changes in magma plumbing systems.

271 Geodetic measurements provide information only about the parts of a plumbing system that  
272 are currently active, and do not necessarily reflect the full extent and character of the intrusion  
273 network (e.g., Sigmundsson, 2016). However, geodetic analyses of ground deformation provide  
274 critical insight into the spatial and temporal development of active plumbing systems. Comparing  
275 observations of ancient plumbing systems (e.g., Magee *et al.*, 2013; Schofield *et al.*, 2014),  
276 integration of ground deformation measurements with petrological observations (e.g., Caricchi *et*  
277 *al.*, 2014; Jay *et al.*, 2014) or thermal models (Parker *et al.*, 2016), as well as tomographic  
278 geophysical imaging, will increase the sophistication of models of magmatic systems. Integrating  
279 InSAR with gravity or electromagnetic measurements is particularly powerful, as it can allow  
280 discrimination between melt, volatiles, and hydrothermal fluids for which deformation signals are  
281 similar (see section 2.4) (e.g., Tizzani *et al.*, 2009).

282

## 283 **2.2. Seismicity and magma plumbing systems**

### 284 ***Technique***

285 Seismicity (i.e. earthquakes) at volcanoes is primarily caused by the dynamic interaction of magma  
286 and hydrothermal fluids with the solid host rock (e.g., Chouet & Matoza, 2013), as well as by  
287 fracturing and fragmentation of silicic magma (e.g., Tuffen *et al.*, 2008). There are a number of

288 primary physical mechanisms for causing volcano seismicity (e.g., faulting), each of which  
289 typically produces seismic signals of specific frequency content (Chouet & Matoza, 2013).  
290 Recording and isolating different volcano seismicity signals therefore allows a variety of plumbing  
291 system processes to be assessed. The majority of volcano monitoring agencies have now deployed  
292 or aim to use a network of distributed seismic sensors, including broadband seismometers, to  
293 monitor volcano activity (Neuberg *et al.*, 1998; Sparks *et al.*, 2012). Furthermore, an increase in  
294 computing power and reduction in cost of seismic sensors means that researchers are now  
295 developing fast, fully automated detection and real-time location techniques that can locate  
296 seismicity to sub-decimetre precision (e.g., Drew *et al.*, 2013; Sigmundsson *et al.*, 2015).

297

## 298 ***Observations***

299 Volcano-tectonic (VT) seismicity generally produces relatively high frequency (1–20 Hz), short  
300 period signals, involving clear primary (P), secondary (S), and surface waves, which are caused by  
301 displacement on new or existing faults in the host rock in response to fluid-induced stress changes  
302 (e.g., Rubin & Gillard, 1998; Roman & Cashman, 2006; Tolstoy *et al.*, 2008). These earthquakes  
303 commonly occur near the propagating edge of intrusions, meaning the space-time evolution of VT  
304 earthquake locations can be used to track the horizontal and vertical growth of sills and dykes (e.g.,  
305 Keir *et al.*, 2009; Sigmundsson *et al.*, 2010; Sigmundsson *et al.*, 2015). Inflation of a magma or  
306 mush body can also induce VT seismicity on any preferentially oriented faults surrounding the  
307 intrusion, thereby recording the delivery time and locus of new magma injected into a reservoir  
308 (e.g., Roman & Cashman, 2006; Vargas-Bracamontes & Neuberg, 2012).

309 Earthquakes with longer period seismic signals and low-frequencies (0.5–2 Hz) are thought  
310 to be generated near the interface between magma and solid rock (Chouet & Matoza, 2013). The  
311 earthquake source proximity to the magma causes the seismic signal to resonate in parts of the  
312 plumbing system (e.g., conduits, dykes, and cracks), leading to a reduction in its frequency content  
313 (Chouet & Matoza, 2013). These earthquakes can potentially be caused by stick-slip motion

314 between the magma and wall-rock or fracturing of cooling magma near the conduit wall (Neuberg  
315 *et al.*, 2006; Tuffen *et al.*, 2008). Such earthquakes typically occur at restricted portions of conduits  
316 where the magma flow and shear strain rate are highest (Neuberg *et al.*, 2006; Tuffen *et al.*, 2008).

317 Very long period seismicity (VLP) of 10s of seconds to several minutes period are typically  
318 attributed to inertial forces associated with perturbations in the flow of magma and gases through  
319 conduits (Chouet & Matoza, 2013). These signals can record the response of the host rock to  
320 reservoir inflation and deflation and may be used to model conduit shape and size (Chouet *et al.*,  
321 2008). To do this requires a better understanding of the links between flow processes and resultant  
322 pressure/momentum changes using laboratory experiments and numerical models that include the  
323 elastic response to magma flow across multiple signal frequency bands (e.g., Thomas & Neuberg,  
324 2012).

325

### 326 ***Implications and integration***

327 Studies of evolving reservoirs now aim to link episodes of seismicity related to new magma  
328 injection to petrological evidence for timing of reservoir recharge events, thereby providing  
329 independent constraints on day to year-long time-scales of magma residence and input prior to  
330 eruptions. For example, Fe-Mg diffusion chronometry modelling of orthopyroxene crystals from the  
331 1980–1986 eruption of Mount St. Helens indicates that compositionally distinct rims grew within  
332 12 months prior to eruption (Fig. 4) (Saunders *et al.*, 2012). Peaks in crystal growth correlated  
333 extremely well with increased seismicity and SO<sub>2</sub> flux (Fig. 4), confirming the relationship between  
334 seismicity and magma movement, as well as demonstrating how a combination of seismicity and  
335 petrological information can be used to detect magma injections (Saunders *et al.*, 2012).

336 Petrology and seismicity can also be integrated with other methods, such as GPS and  
337 InSAR. Field *et al.*, (2012) analysed volatiles in melt inclusions trapped in phenocrysts within  
338 peralkaline lavas from historic eruptions at the Dabbahu Volcano in Afar, Ethiopia. Volatile  
339 saturation pressures at typical magmatic temperatures were constrained to be in the range 43–207

MPa, consistent with the phenocryst assemblage being stable at 100–150 MPa. The interpreted magma/mush storage depths for these historic eruptions are ~1–5 km, consistent with the depths of earthquakes associated with reservoir inflation following dyke intrusion in 2005–2006 (Fig. 5) (Ebinger *et al.*, 2008; Field *et al.*, 2012). Additionally, the best-fit result for modelling of uplift patterns recorded by InSAR data, which were collected over the same time period as seismicity measurement, suggests the magma/mush reservoir comprises a series of stacked sills over a ~1–5 km depth range (Fig. 5) (Ebinger *et al.*, 2008). The consistency of depth estimates based on petrological study of ancient eruptions, along with the seismicity and inflation of the Dabbahu Volcano following axial dyke intrusion in 2005–2006, implies a vertically extensive and potentially long-lived magma/mush storage region. Such multidisciplinary studies demonstrate that joint observations and modelling of seismic signals, petrological data, and other techniques (e.g., geodesy and gas emissions) significantly strengthen interpretation of the physical structure, emplacement, and evolution of magma plumbing systems.

### 2.3. Identifying melt in plumbing systems using seismic imaging

#### *Techniques*

Both active and passive source seismological techniques, which utilise man-made seismic events and natural earthquakes respectively, can be used to identify areas where the presence of partial melt or magma causes a local reduction in seismic wavespeed, an increase in anisotropy, or an increase in attenuation (e.g., Berryman, 1980; Hammond & Humphreys, 2000a, b). With the recent availability of dense seismic networks, resolution of the crust and mantle seismic velocity structure has improved to the degree that active source seismic experiments can: (i) use tomographic techniques to image likely storage regions in the upper crust beneath ocean island volcanoes (e.g., Soufrière Hills Volcano, Montserrat; Fig. 6) (Paulatto *et al.*, 2010; Shalev *et al.*, 2010) and, occasionally, onshore volcanoes (e.g., Mt Erebus, Antarctica, Zandomenighi *et al.*, 2013; Mt. St. Helens, Kiser *et al.*, 2014); and (ii) utilise reflected data to image individual sills beneath mid-ocean

366 ridges (e.g., Kent *et al.*, 2000, Marjanovic *et al.*, 2014). A further example from Katla volcano  
367 Iceland, demonstrates how active source seismic experiments can be used to identify S-wave  
368 shadow zones (i.e. S-waves cannot travel through fluids) and delays in P-waves, which may be used  
369 to infer the location and geometry of shallow-level magma reservoirs (Gudmundsson *et al.*, 1994).  
370 However, recent modelling approaches suggest that the upper crust likely represents only a small  
371 portion of magma plumbing systems and long-term storage is dominated by mushy zones  
372 throughout the lower crust (e.g., Annen *et al.*, 2006). Active source seismic experiments,  
373 particularly on land where the crust is thick and coverage less uniform, cannot penetrate to these  
374 depths efficiently. Furthermore, whilst seismic tomographic methods using local earthquakes offer  
375 3D images of crustal velocity beneath many volcanoes (e.g., Mt. St. Helens, Waite & Moran, 2009;  
376 Askja, Iceland, Mitchell *et al.*, 2013), they can only resolve areas directly above the deepest  
377 earthquakes. Non-uniform coverage thus makes interpreting tomographic images difficult as  
378 resolution varies across the model (see review by Lees, 2007).

379         To illuminate lower crustal regions, seismologists rely on passive seismology. Extending  
380 seismic tomographic images of magma plumbing systems to lower crustal depths requires the use of  
381 teleseismic body-wave and surface wave data, which emanate far (>1000 km) from the  
382 measurement site. However, these data are dominated by longer period signals, meaning their  
383 resolution is relatively low. For example, the Fresnel zone (i.e. the region within  $\frac{1}{4}$  seismic  
384 wavelength and an estimate of the minimum resolvable structure) for active source data at 10 Hz is  
385 on the order of 3 km in the upper crust compared to 10–15 km for 1 Hz teleseismic data used in  
386 receiver function or tomography studies.

387

## 388 ***Observations***

389 Active and passive seismological techniques provide crucial insight into transcrustal melt and  
390 magma distribution. For example, P-wave seismic travel-time tomography across Monserrat and the  
391 Soufrière Hills Volcano images a series of relatively fast seismic velocity zones, which are



392 interpreted as solidified andesitic intrusions, surrounded by regions of slow seismic velocities likely  
393 related to either areas of hydrothermal alteration or buried volcanoclastic deposits (Fig. 6) (Paulatto  
394 *et al.*, 2010; Shalev *et al.*, 2010). Within the lower crust, inversions using surface wave data  
395 generated by ambient seismic noise and receiver function data, which isolates P-wave to S-wave  
396 conversions at major discontinuities in the earth, have identified low shear-wave velocities probably  
397 related to melt presence beneath several volcanic settings (e.g., New Zealand, Bannister *et al.*, 2007;  
398 Toba, Sumatra, Stankiewicz *et al.*, 2010; Ethiopia, Hammond *et al.*, 2011; Jaxybulatov *et al.*, 2014;  
399 Costa Rica, Harmon & Rychert, 2015).

400         When trying to determine how much melt or magma is present, numerous studies have  
401 shown that seismic velocities are more sensitive to the shapes of melt/magma-filled spaces on a  
402 range of scales compared to the melt fraction (e.g., Hammond & Humphreys, 2000a, b; Miller &  
403 Savage, 2001; Johnson & Poland, 2013; Hammond & Kendall, 2016). On the grain-scale, melt  
404 commonly wets grain boundaries, forming planar pockets (e.g., Takei, 2002; Garapic *et al.*, 2013;  
405 Miller *et al.*, 2014), whereas on the larger scale magma may form planar intrusions of either mush  
406 (e.g., Annen *et al.*, 2006), or liquid-rich dykes or sills. If these features are preferentially aligned,  
407 they will appear as a distributed region of melt to seismic waves and the analyses described will not  
408 be able to discriminate between a melt-poor region dominated by aligned melt-pockets on grain  
409 boundaries and an elongate melt-rich body such as an intrusion (e.g., Hammond & Kendall, 2016).  
410 A further problem is that seismic velocities are affected by variations in temperature (Jackson *et al.*,  
411 2002), composition (Karato & Jung, 1998), and attenuation (Goes *et al.*, 2012). Relating seismic  
412 velocity anomalies to melt fraction is therefore difficult without some prior knowledge of melt  
413 distribution (Hammond & Kendall, 2016).

414         One possible approach to investigate melt distributions further is through measuring seismic  
415 anisotropy. If melt has some preferential distribution on a length-scale smaller than the seismic  
416 wavelength, such as a stacked network of sills or an anisotropic permeability on the grain scale,  
417 then the seismic wavespeed will vary with direction of propagation, i.e. be anisotropic. As a result,

418 measuring the effects of seismic anisotropy allows inferences about sub-seismic wavelength  
419 structures, leading to estimates of the preferential orientation of melt distribution. It is common to  
420 observe strong anisotropy beneath volcanoes and this has been used to place constraints on melt  
421 distribution. For example, high degrees of shear-wave splitting from volcanic earthquakes can either  
422 directly map out regions of significant quantities of melt aligned in pockets (Keir *et al.*, 2011), or  
423 map out stress changes related to overpressure from injections of magma into the upper crust (Gerst  
424 & Savage, 2004; Roman *et al.*, 2011). To image the deeper crustal magmatic system, azimuthal  
425 variations in the ratio of P-wave to S-wave speeds (i.e.  $V_p/V_s$ ) from receiver functions led to the  
426 interpretation that a stacked network of sills is present in the lower crust beneath the Afar  
427 Depression, Ethiopia (Hammond, 2014). Differences in the velocity of Rayleigh Waves and Love  
428 Waves, which are vertically polarised shear-waves and horizontally polarised shear waves  
429 respectively, suggest a similar anisotropic melt distribution is present beneath the Toba Caldera,  
430 Sumatra (Jaxybulatov *et al.*, 2014) and Costa Rica (Harmon & Rychert, 2015).

431

### 432 ***Implications and integration***

433 Due to the interference of signals denoting the geometry of melt-filled pockets and the volumetric  
434 proportion of that melt, estimating melt fraction remains difficult using seismology alone. Some  
435 attempt has been made to directly infer magma/mush reservoir properties from seismic velocities.  
436 For example, Paulatto *et al.*, (2012) used thermal modelling to test the range of melt fractions that  
437 could account for the low velocity zones imaged in the upper crust beneath Soufrière Hills Volcano  
438 (Fig. 6), Montserrat and concluded the melt fraction is between 3 and 10%. However, accounting  
439 for resolution of the tomography, together with uncertainties in the distribution and geometry of  
440 melt, means >30% melt may be present more locally in the low velocity zones defined beneath  
441 Soufrière Hills Volcano (Paulatto *et al.*, 2012). Possible ways forward involve integrating  
442 seismological data with: (i) petrological data that can place limits on likely melt fractions and/or  
443 emplacement depths (e.g., McKenzie & O’Nions, 1991; Comeau *et al.*, 2016); (ii) geochemical

444 techniques that can help determine timescales of melt and magma evolution (e.g., Hawkesworth *et*  
445 *al.*, 2000); and (iii) geodetic or other monitoring data, which helps determine magma movement  
446 (Sturkell *et al.*, 2006). Recent efforts applying industry software, such as full waveform inversions  
447 (FWI; Warner *et al.*, 2013), which is discussed in section 3.1, are also pushing the potential  
448 application of seismological data further and mean that it may be possible to resolve features to sub-  
449 kilometre levels, particularly in the upper crust. Together, these techniques may allow us to directly  
450 relate seismic velocity anomalies to melt fractions and distributions in the whole crust.

451

## 452 **2.4. Studying magma plumbing systems using gravimetry**

### 453 ***Techniques***

454 Gravimetry measures the gravitational field and its changes over space and time, which can be  
455 related to variations in the subsurface distribution and redistribution of mass (e.g., magma). A  
456 variety of gravimeter instruments (e.g., free-fall, superconducting, and spring-based) and techniques  
457 (e.g., ground-based, sea-floor, ship-borne, and air-borne instrumentations) are available. Spring  
458 gravimeters, where a test mass is suspended on a spring, are mostly used to study magmatic and  
459 volcanic processes in ground-based surveys (e.g., Carbone *et al.*, 2017; Van Camp *et al.*, 2017).  
460 Changes in the gravitational acceleration across a survey area shorten or lengthen the spring, which  
461 is recorded electronically and converted to gravity units. These changes are evaluated across a  
462 survey network in relation to a reference and are hence termed ‘relative measurements’. Absolute  
463 gravimetry can also be measured, i.e. the value of gravitational acceleration, and serves primarily to  
464 create a reference frame into which other geodetic methods (e.g., InSAR, GNSS, levelling, relative  
465 gravimetry) can be integrated for joint data evaluation. Recent reviews by Carbone *et al.*, (2017)  
466 and Van Camp *et al.*, (2017) provide a broad account of gravimetric instruments, measurement  
467 protocols, and data processing relevant for the study of magmatic systems.

468         Static gravimetric techniques obtain a single snap-shot of the subsurface mass distribution.

469         For example, Bouguer anomaly maps are perhaps the best-known products of static gravity surveys

470 and capture spatial variations in gravity over an area of interest, providing insight into anomalous  
471 mass distribution in the subsurface. Within magmatic studies, computational modelling and  
472 inversion of Bouger anomaly data allows identification of shallow intrusions (e.g., dykes and sills;  
473 Rocchi *et al.*, 2007), magma-related ore bodies (Hammer, 1945; Bersi *et al.*, 2016), and plutons  
474 (e.g., Figs 7a and b) (e.g., Vigneresse, 1995; Vigneresse *et al.*, 1999; Petford *et al.*, 2000) exhibiting  
475 a density contrast with their host rocks.

476 In contrast to static surveys, dynamic gravimetric observations allow spatio-temporal mass  
477 changes to be tracked. Dynamic gravimetric studies investigate how the subsurface architecture  
478 changes over time and is usually performed by measuring variations in gravity across a network of  
479 survey points (e.g., Fig. 7c) or, in a few exceptional cases, by installing a network of continuously  
480 operating gravimeters. Dynamic observations demand one-to-two orders of magnitude higher data  
481 precision (i.e. to a few  $\mu\text{Gal}$  where  $1 \mu\text{Gal} = 10^{-8} \text{ m/s}^2$ ) compared to static surveys, making them an  
482 elaborate and time-consuming exercise. However, dynamic gravity data yields important insights  
483 into the source processes behind non-tectonic volcano and crustal deformation, particularly if  
484 combined with surface deformation data (e.g., InSAR and GNSS) as subsurface mass and volume  
485 changes can be employed to characterise the density of the material behind the stress changes (Figs  
486 7c-f and 8) (e.g., Battaglia & Segall, 2004; Jachens & Roberts, 1985; Poland & Carbone, 2016).  
487 There are also cases where volcano unrest, due either to magma intrusion into a ductile host rock or  
488 to volatile migration at shallow depths, does not result in resolvable surface deformation; in these  
489 scenarios, gravity data have provided vital clues about subsurface processes otherwise hidden from  
490 conventional monitoring techniques (e.g., Gottsmann *et al.*, 2006; Gottsmann *et al.*, 2007; Miller *et*  
491 *al.*, 2017).

492 Whilst static and dynamic gravimetric observations offer considerable insight into the  
493 structure and dynamics of magma plumbing systems, care must be exercised when collecting and  
494 interpreting gravity data from active magmatic areas where seasonal variations in hydrothermal  
495 systems, aquifers, or the vadose zone can influence subsurface mass distribution (e.g., Hemmings *et*

496 *al.*, 2016). These seasonal changes can, in some cases, result in data aliasing artefacts and inhibit  
497 the quantification of deeper-seated magmatic processes (e.g., Gottsmann *et al.*, 2005; Gottsmann *et*  
498 *al.*, 2007).

499

## 500 ***Observations***

501 Gravimetric investigations have been at the heart of studies into the subsurface structure of active  
502 and ancient magma plumbing systems for more than 80 years (e.g., Carbone *et al.*, 2017; Van Camp  
503 *et al.*, 2017). Using techniques initially designed for imaging salt domes, silicic plutons were the  
504 first components of magma plumbing systems to be examined using gravimetry because their low  
505 density relative to surrounding rocks produces clear, negative gravity anomalies of ~10 to ~40 mGal  
506 amplitude (e.g., Reich, 1932; Bucher, 1944; Bott, 1953). Gravity data have been instrumental in the  
507 investigation of upper-crustal, silicic magma plumbing systems, helping to reveal: (i) the 3D  
508 geometry of plutons by allowing floor morphologies (e.g., flat-floored or wedge-shaped) to be  
509 determined (e.g., Vigneresse *et al.*, 1999; Petford *et al.*, 2000); and (ii) how plutons are constructed,  
510 for example, by the amalgamation of multiple intrusions fed from depth by dykes (e.g., Vigneresse,  
511 1995). Furthermore, recent high-precision static surveys over active silicic volcanoes have enabled  
512 detailed modelling of the sub-volcanic magma plumbing system, commonly demonstrating the  
513 occurrence of vertically extensive, transcrustal magma bodies (Figs 7a and b) (e.g., Gottsmann *et*  
514 *al.*, 2008; del Potro *et al.*, 2013; Saxby *et al.*, 2016; Miller *et al.*, 2017). In addition to examining  
515 silicic magma plumbing systems, negative gravity anomalies with typical amplitudes of up to 60  
516 mGal and up to 100 km wavelength can be associated with, and provide insight into, the geometry  
517 and size of silicic ash-flow calderas (e.g., Eaton *et al.*, 1975; Masturyono *et al.*, 2001). Positive  
518 gravity anomalies with amplitudes of up to 30 mGal and wavelengths of up to 20 km are  
519 commonly identified at mafic volcanoes and likely result from dense intrusive complexes (e.g.,  
520 Rymer & Brown, 1986).

Dynamic gravity observations have provided unprecedented insight into the evolution of magma plumbing systems over timescales of seconds to decades, including: (i) the characterisation of multi-year lava lake dynamics (e.g., Poland & Carbone, 2016); (ii) mass budgets of magma intrusions (e.g., Fig. 8) (e.g., Battaglia *et al.*, 1999; Jousset *et al.*, 2000; Rymer *et al.*, 2005; Bonforte *et al.*, 2007; Tizzani *et al.*, 2009); (iii) shallow hydrothermal fluid flow processes induced by deeper magmatic unrest (e.g., Battaglia *et al.*, 2006; Gottsmann *et al.*, 2007; Miller *et al.*, 2017); and (iv) parameters of magmatic geothermal reservoirs (e.g., Hunt & Bowyer, 2007; Sofyan *et al.*, 2011). For example, using data from a network of continuously recording gravimeters, Carbone *et al.*, (2013) calculated the density of the Kilauea lava lake as  $950 \pm 300 \text{ kg m}^{-3}$ , i.e. similar to and potentially less than that of water, suggesting that the magma column within the upper portions of the volcanic edifice is gas-rich. Because density and volatile content are critical controls on magma rheology, identification of a gas-rich magma column and lava lake at Kilauea is crucial to modelling and understanding convection and eruption dynamics (Carbone *et al.*, 2013).

### ***Implications and integration***

The advent of data-rich geodetic observations from satellite-remote sensing (e.g., InSAR), in conjunction with spatio-temporal gravity studies, provides unprecedented opportunities to characterise magma plumbing system dynamics and the driving mechanisms behind volcano deformation. At Long Valley caldera, for example, a residual gravity increase of more than  $60 \text{ } \mu\text{Gal}$  between 1982 and 1999 indicates a mass addition at depth (Battaglia *et al.*, 1999). Joint inversion of InSAR and gravity data from Long Valley derives a best fit-source density of  $2509 \text{ kg m}^{-3}$  and is indicative of a magmatic intrusion (Fig. 8) (Tizzani *et al.*, 2009). At the deforming Laguna del Maule volcanic centre, Chile, multi-year InSAR and dynamic gravity records demonstrate that uplift and extension above an inflating sill-like reservoir at  $\sim 5 \text{ km}$  depth promoted migration of hydrothermal fluids along a fault to shallow ( $1\text{--}2 \text{ km}$ ) depths (Miller *et al.*, 2017). Alternatively, although no ground deformation is observed at Tenerife, Spain, deconvolution of dynamic gravity

547 into a shallow and deep gravity field provides evidence of unrest (Prutkin *et al.*, 2014). The gravity  
548 data suggest hybrid processes have generated the unrest, whereby fluids were released and migrated  
549 upward along deep-rooted faults from an intrusion at ~9 km beneath the summit of Teide Volcano  
550 (Prutkin *et al.*, 2014). Overall, combining ground deformation and gravimetric observations has  
551 highlighted complex processes both within magma reservoirs (e.g., mass addition by magma input,  
552 density decrease by volatile exsolution, or density increase by crystallisation; Figs 7c-f) and in the  
553 surrounding host rock (e.g., migration of magmatic fluids, phase changes in hydrothermal systems).  
554 Key to a better understanding of the processes governing these magma plumbing system and  
555 volcano deformation dynamics is the integration of gravimetric and geodetic data with other  
556 geophysical data (e.g., seismicity or magnetotellurics) and petrological data. Coupled with  
557 advanced numerical modelling, such multi-parameter studies promise exciting new insights into the  
558 inner workings of sub-volcanic magma plumbing systems (e.g., Currenti *et al.*, 2007; Hickey *et al.*,  
559 2016; Currenti *et al.*, 2017; Gottsmann *et al.*, 2017; Miller *et al.*, 2017).

560

## 561 **2.5. Resolving magma plumbing system structure with electromagnetic methods**

### 562 ***Techniques***

563 Electromagnetic (EM) methods probe subsurface electrical resistivity or its inverse, i.e. electrical  
564 conductivity. Spatial variations in resistivity control the position, strength, and geometry of local  
565 electrical eddy currents and the magnetic fields they produce. These electrical eddy currents are  
566 induced by time-varying, naturally occurring magnetic fields external to Earth, which forms the  
567 basis of the magnetotelluric (MT) technique, or by controlled sources. Monitoring these decaying  
568 electrical and magnetic fields with passive MT techniques therefore allows the subsurface resistivity  
569 distribution to be inferred. Controlled source methods generally probe only the shallow subsurface,  
570 but MT has a greater depth range as it uses longer-period signals to penetrate deeper. The signals  
571 propagate diffusively, which means EM methods typically have a lower resolution than seismic  
572 techniques. However, melt, magma, and magmatic hydrothermal fluids are generally considerably

573 less resistive than solid rock and can thus easily be detected by EM methods, which are sensitive to  
574 conductive materials (e.g., Whaler & Hautot, 2006; Wannamaker *et al.*, 2008; Desissa *et al.*, 2013;  
575 Comeau *et al.*, 2015). EM methods, particularly MT, have therefore been used extensively to study  
576 magmatic systems in various tectonic settings.

577 MT equipment, data acquisition, and processing is described by Simpson & Bahr (2005) and  
578 Ferguson (2012). Measured field variations have very low amplitudes, meaning equipment needs to  
579 be positioned and installed carefully to reduce vibrational (e.g., from wind, vegetation, or vehicles)  
580 and electrical (e.g., from power lines) noise. If data are recorded synchronously at a second, less  
581 noisy site, remote reference methods can be used to improve the data quality (e.g., Gamble *et al.*,  
582 1979). One further problem is that small-scale resistivity anomalies in the shallow subsurface  
583 generate galvanic (non-inductive) effects that distort MT data. The distortion is identified and  
584 corrected for, which may involve using controlled source transient electromagnetic data to ensure  
585 complete removal (e.g. Sternberg *et al.*, 1988), at the same time as assessing whether the data can  
586 be modelled with a one-, two- or three-dimensional resistivity structure (e.g. Jones, 2012). Failure  
587 to remove galvanic distortion can result in models having resistivity features at the wrong depth.  
588 For example, there has been controversy as to whether a conductor beneath Vesuvius Volcano, Italy  
589 is caused by a deep (~8–10 km depth) magma reservoir (Di Maio *et al.*, 1998) or a shallow brine  
590 layer (Manzella *et al.*, 2004). All of these factors can be a significant problem when using MT to  
591 study magmatic systems, especially on volcanic islands.

592 The relationship between MT data and subsurface resistivity is strongly non-linear meaning  
593 that inversion is fundamentally non-unique and computationally expensive (e.g., Bailey, 1970;  
594 Parker, 1980; Weaver, 1994). Most practical algorithms for inverting MT data obtain a unique  
595 result by minimising a combination of misfit to the data and a measure of model roughness (e.g.,  
596 Constable *et al.*, 1987). This approach poorly delimits how magma is distributed in the subsurface,  
597 whether it is in sills, dykes, or larger reservoirs (Johnson *et al.*, 2016). Whilst MT data are sensitive  
598 to the top surface of a conductor, its base may not be detected because conductive material reduces



599 the penetration depth of the signal. Sensitivity analysis is used to ascertain the model features  
600 required to fit the MT data, which allows a conductor to be confined to a certain depth range and  
601 thereby constrains its base (e.g., Desissa *et al.*, 2013). Furthermore, if the resistivity of a conductor  
602 can be inferred, its conductance (i.e. a product of thickness and conductivity) can be used to  
603 determine its thickness (e.g., Comeau *et al.*, 2016).

604

## 605 ***Observations***

606 EM induction surveys have been conducted on most major sub-aerial volcanoes and magmatic  
607 systems; only a few will be mentioned here to illustrate the type information on magma plumbing  
608 systems that has been obtained. MT data have been used to image several low resistivity features in  
609 the central Andes, particularly beneath the uplifting (10–15 mm/yr) Volcán Uturuncu, Bolivia (Fig.  
610 9a) (Comeau *et al.*, 2015; Comeau *et al.*, 2016). The deepest of these bodies has resistivities of  $<3$   
611  $\Omega$  m, has a top contact at ~15–20 km depth (i.e. it is shallowest beneath Uturuncu), likely has a  
612 thickness of  $>6$  km, and extends E-W for ~170 km (Fig. 9) (Comeau *et al.*, 2015; Comeau *et al.*,  
613 2016). This large-scale structure is interpreted to be the Altiplano-Puna magma body (APMB),  
614 which has been identified in other geophysical datasets (e.g., Fig. 7a) (e.g., gravimetry, del Potro *et*  
615 *al.*, 2013), with its low resistivity attributed to the presence of at least 20% andesitic melt and/or  
616 magma. Extending from the top of the APMB towards the surface are several vertical, narrow ( $<10$   
617 km wide), low resistivity ( $<10 \Omega$  m) zones that coincide with areas of seismicity and negative  
618 gravity anomalies (Fig. 9). These zones likely reflect a network of dykes and upper crustal magma  
619 reservoirs (Jay *et al.*, 2012; del Potro *et al.*, 2013; Comeau *et al.*, 2015; Comeau *et al.*, 2016).

620 Monitoring of magmatic systems can also be undertaken by both time-lapse and continuous  
621 EM measurement. For example, MT data collected immediately after the 1977–1978 eruption at  
622 Usu volcano, Japan revealed a conductive zone ( $<100 \Omega$  m) beneath the summit that probably  
623 corresponded to intruded magma. By 2000, MT data revealed that this conductive body had become  
624 resistive (500–1000  $\Omega$  m) as the intrusion cooled, from 800°C to 50°C, and crystallised

625 (Matsushima *et al.*, 2001). Continuous MT monitoring of Sakurajima volcano, Japan between May  
626 2008 and July 2009 revealed temporal changes in resistivity of  $\pm 20\%$ , some of which correlated to  
627 periods of surface deformation and were inferred to reflect mixing between groundwater and  
628 volatiles exsolved from an underlying magma body (Aizawa *et al.*, 2011). Continuous MT  
629 monitoring at La Fournaise, Réunion Island recorded apparent resistivity decreases associated with  
630 the large 1998 eruption, which were attributed to the injection of a N-S striking dyke (Wawrzyniak  
631 *et al.*, 2017).

632         Several EM studies have focussed on magma plumbing systems at divergent margins,  
633 including mid-ocean ridges and continental rifts. For example, at the fast-spreading East Pacific  
634 Rise, a ~10 km wide, sub-vertical conductor, slightly displaced from the ridge axis and connected to  
635 a deep, broad conductive zone was interpreted as a channel efficiently transporting melt to the base  
636 of the crust (Baba *et al.*, 2006; Key *et al.*, 2013). Imaging of a crustal conductor for the first time  
637 beneath a slow-spreading ridge, i.e. the Reykjanes ridge in the Atlantic Ocean, suggests that magma  
638 injection into crustal reservoirs is intermittent but rapid (MacGregor *et al.*, 1998; Heinson *et al.*,  
639 2000). Conversely, slow-spreading continental rifting in the Dabbahu magma segment, Afar,  
640 Ethiopia appears to be underlain by a large conductor, either at the top of the mantle or straddling  
641 the Moho, containing more melt ( $>300 \text{ km}^3$ ) than is intruded into the magma plumbing system  
642 during a typical rifting episode (Desissa *et al.*, 2013). The volume of this large conductor implies it  
643 is a long-lived feature that could source magmatic activity for tens of thousands of years (Desissa *et*  
644 *al.*, 2013).

645

## 646 ***Implications and integration***

647 It is clear from MT studies of the APMB that other geophysical techniques aid and/or corroborate  
648 data interpretation (Fig. 9) (e.g., Comeau *et al.*, 2015; Comeau *et al.*, 2016). Over the last two  
649 decades, numerous geophysical studies have been applied to examine magma and melt distribution  
650 beneath various portions of the East African Rift, providing an excellent opportunity to test how

651 different techniques and data can be integrated. For example, extensive zones of melt beneath the  
652 Afar region in Ethiopia inferred from MT data by Desissa *et al.*, (2013) is supported by: (i) the  
653 occurrence of coincident, low P-wave velocity (down to  $7.2 \text{ km s}^{-1}$ ) zones identified using from  
654 analysis of seismic Pn waves that propagate along the Moho (Stork *et al.*, 2013); (ii) surface wave  
655 studies that reveal lower crustal areas in magmatic domains with low S-wave velocities ( $\sim 3.2 \text{ km s}^{-1}$ ) (Guidarelli *et al.*, 2011); and (iii) high anisotropic  $V_p/V_s$  ratios and low amplitude receiver  
656 functions, which are indicative of the presence of melt (Hammond *et al.*, 2011; Hammond, 2014).  
657 Similarly, crustal conductors along the northern flanks of the Main Ethiopian Rift, interpreted to  
658 represent melt/magma (Whaler & Hautot, 2006; Samrock *et al.*, 2015; Hübner *et al.*, 2018), coincide  
659 with locations where receiver functions either have amplitudes too low to interpret or indicate high  
660  $V_p/V_s$  values (Dugda *et al.*, 2005; Stuart *et al.*, 2006). Electrical anisotropy can be inferred directly  
661 from MT data consistent with a two-dimensional subsurface resistivity distribution (Padilha *et al.*,  
662 2006; Hamilton *et al.*, 2006). Large amounts of electrical anisotropy were found in the lower crust  
663 beneath Quaternary magmatic segments in Afar, Ethiopia, where there is also significant crustal  
664 seismic anisotropy (see Fig. 11 of Ebinger *et al.*, 2017); oriented melt-filled pockets are the  
665 probable cause of both.

667 Although EM methods can image subsurface conductors that are interpreted to represent  
668 magma bodies or zones of partial melt (i.e. crystal mushes), additional information is required to  
669 determine their composition, volume, and/or melt fraction. However, there are several challenges in  
670 inverting measured bulk resistivities to recover this information. Two-phase mixing laws predict  
671 bulk resistivity is primarily a function of melt resistivity and geometry in the rock matrix when the  
672 fluid phase has low resistivity, as in the case of partial melt. Well-connected melt gives a lower bulk  
673 resistivity than isolated melt pockets, for the same melt fraction and resistivity (e.g. Hashin &  
674 Shtrikman, 1963; Roberts & Tyburczy, 1999; Schmeling, 1986). Whilst resistivities of basaltic and  
675 rhyolitic melts have been measured in laboratory experiments (e.g., Laumonier *et al.*, 2015; Guo *et al.*, 2016), they are strongly dependent on temperature, pressure, silica, sodium and water content,

677 making extrapolation uncertain. The web-based SIGMELTS tool can, however, be used to predict  
678 melt and bulk resistivities for a wide range of compositions and conditions (Pommier & Le Trong,  
679 2011). Importantly, petrological and geochemical characterisation of eruptive products can help  
680 inform interpretations of associated, subsurface conductors but it is difficult to ascertain either  
681 whether their composition reflects the current magma/melt present in the plumbing system or  
682 whether melt pockets are interconnected. These large uncertainties in melt resistivity and the  
683 requirement to make assumptions about its geometry make direct inference of melt fraction  
684 difficult. Nonetheless, information from laboratory studies, petrology, and geochemistry aids  
685 interpreting resistivity anomalies in magmatic regions (see review by Pommier, 2014).

686

## 687 **2.6. Imaging ancient magma plumbing systems in seismic reflection data**

### 688 ***Techniques***

689 Over the last two decades, major advances have been made in imaging deep crustal melt beneath  
690 active volcanic terrains using P- and S-wave tomographic data (e.g., Yellowstone, Husen *et al.*,  
691 2004; Mt. St. Helens, Lees, 2007; Hawaii, Okubo *et al.*, 1997). These data image deep (>7 km),  
692 often laterally extensive (up to 20 km), sill-like magma reservoirs (e.g., Paulatto *et al.*, 2012).  
693 However, like many geophysical and geodetic techniques applied to study active magma plumbing  
694 systems, these data typically lack the spatial resolution to resolve the detailed geometry of pathways  
695 transporting magma to the Earth's surface. Active source seismic reflection data, which have a  
696 spatial resolution of metres-to-decametres down to depths of ~5 km, can provide unprecedented  
697 images of and insights into the geometry and dynamics of shallow-level, crystallised, magma  
698 plumbing systems (e.g., Fig. 10) (e.g., Planke *et al.*, 2000; Smallwood & Maresh, 2002; Thomson &  
699 Hutton, 2004; Cartwright & Hansen, 2006; Jackson *et al.*, 2013; Magee *et al.*, 2016; Schofield *et*  
700 *al.*, 2017). Whilst seismic reflection data are traditionally used to find and assist in the production of  
701 hydrocarbons in sedimentary basins (Cartwright & Huuse, 2005), we here discuss and support its  
702 application to volcanological problems.

703           Acquiring active source seismic reflection data involves firing acoustic energy (i.e. seismic  
704 waves) into the subsurface and measuring the surface arrival times (i.e. the travel-time) of reflected  
705 energy. Processing of these arrival time data allows reconstruction of the location and geometry of  
706 the geological interfaces from which acoustic energy was reflected. Mafic intrusive igneous rocks  
707 are generally well-imaged in seismic reflection data because they typically have greater densities  
708 ( $>2.5 \text{ g/cm}^3$ ) and acoustic velocities (i.e.  $>4000 \text{ m/s}$ ) than encasing sedimentary strata; these  
709 differences result in a high acoustic impedance contrast, causing more seismic energy to be  
710 reflected back to the surface compared to low acoustic impedance boundaries (Smallwood &  
711 Maresh, 2002; Brown, 2004). In contrast, silicic igneous rocks have similar acoustic properties to  
712 encasing sedimentary strata, meaning that felsic intrusions are rarely imaged in seismic reflection  
713 data (Mark *et al.*, 2017; Rabbel *et al.*, 2018). Furthermore, because reflection seismology relies on  
714 the return of acoustic energy to the surface, seismic reflection data favourably image mafic, sub-  
715 horizontal-to-moderately inclined intrusions (e.g., sills, inclined sheets, and laccoliths; Smallwood  
716 & Maresh, 2002; Jackson *et al.*, 2013; Magee *et al.*, 2016). Sub-vertical dykes reflect only a limited  
717 amount of acoustic energy back to the surface and are thus typically poorly imaged in seismic  
718 reflection data (e.g., Smallwood & Maresh, 2002; Planke *et al.*, 2005; Thomson, 2007; Wall *et al.*,  
719 2010; Eide *et al.*, 2017a; Phillips *et al.*, 2017).

720

## 721 ***Observations***

722 Sills and inclined sheets are commonly observed in seismic reflection data as laterally  
723 discontinuous, high-amplitude reflections, which may cross-cut the host rock strata (Fig. 10) (e.g.,  
724 Symonds *et al.*, 1998, Smallwood & Maresh, 2002; Planke *et al.*, 2005, Magee *et al.*, 2015). Many  
725 of the sills and inclined sheets imaged in seismic reflection data are, however, expressed as tuned  
726 reflection packages, whereby discrete reflections from the top and base contacts interfere on their  
727 return to the surface and cannot be distinguished (e.g., Figs 10 and 11a) (e.g., Smallwood &  
728 Maresh, 2002; Peron-Pinvidic *et al.*, 2010; Magee *et al.*, 2015; Eide *et al.*, 2017a; Rabbel *et al.*,

2018). It is therefore difficult to assess either intrusion thicknesses, or to detect whether imaged sills are composite bodies made of numerous, stacked, thin sheets. Either way, subtle vertical offsets and corresponding amplitude variations of sill reflections can often be mapped, defining linear structures that radiate out from either the central, deepest portions of sills or areas where underlying intrusions intersect the sill (e.g., Schofield *et al.*, 2012a; Magee *et al.*, 2014; Magee *et al.*, 2016). These structures are interpreted to relate to magma flow indicators such as intrusive steps, broken bridges, and magma fingers (e.g., Schofield *et al.*, 2010; Schofield *et al.*, 2012b; Magee *et al.*, 2018).

A recurring observation from seismic reflection-based studies of extinct and buried intrusive systems is that complexes of interconnected sills and inclined sheets, which may cover  $>3 \times 10^6$  km<sup>2</sup>, can dominate magma plumbing systems (e.g., Fig. 10b) (e.g., Svensen *et al.*, 2012, Magee *et al.*, 2016). Importantly, where buried volcanic edifices are imaged in seismic reflection data, they rarely appear to be underlain by ‘magma chambers’ (i.e. a spheroidal or ellipsoidal body of now-crystallised magma). Instead, these imaged volcanoes commonly appear laterally offset from genetically related sills and/or laccoliths that are inferred to represent their feeder reservoirs (e.g., Fig. 10b) (Magee *et al.*, 2013a; McLean *et al.*, 2017). The geometry, location, and connectivity of these intrusions, which can represent magma storage sites and conduits to the surface, are often heavily influenced by both the host rock structure and lithology (see review by Magee *et al.*, 2016). For example, magma may flow along pronounced discontinuities (e.g., bedding) or within specific stratigraphic units (e.g., coal) for considerable distances, occasionally climbing to higher stratigraphic levels by instigating deformation of the host rock or by exploiting pre-existing faults (e.g., Jackson *et al.*, 2013; Magee *et al.*, 2016; Schofield *et al.*, 2017; Eide *et al.*, 2017b). It is clear from seismic reflection data that shallow-level tabular intrusions are commonly accommodated by roof uplift to form a flat-topped or dome-shaped forced fold (e.g., Figs 11a and b) (e.g., Trude *et al.*, 2003; Hansen & Cartwright, 2006; Jackson *et al.*, 2013; Magee *et al.*, 2013b). Moreover, if the age of reflections onlapping onto these intrusion-induced forced folds can be ascertained, the timing and

755 to some extent the duration of magmatic activity can be determined (e.g., Trude *et al.*, 2003;  
756 Hansen & Cartwright, 2006; Magee *et al.*, 2014; Reeves *et al.*, 2018). Although most seismic-based  
757 studies examine intrusions within sedimentary basins, saucer-shaped sills and laterally extensive  
758 sill-complexes emplaced into crystalline basement rock are also imaged (e.g., Ivanic *et al.*, 2013;  
759 McBride *et al.*, 2018). Lastly, seismic reflection data can also be used to image the internal structure  
760 of layered ultramafic-mafic intrusions (e.g., the Bushveld Layered Intrusion, Malehmir *et al.*, 2012)  
761 and, in some instances, identify dykes (e.g., Fig. 11c) (e.g., Wall *et al.*, 2010; Abdelmalak *et al.*,  
762 2015; Bosworth *et al.*, 2015; Phillips *et al.*, 2017).

763

#### 764 ***Implications and integration***

765 Despite being limited in terms of their spatial resolution (typically a few tens of metres) and ability  
766 to image steeply dipping features (i.e. dykes), they provide unprecedented snapshots into the final  
767 3D structure of magma plumbing systems. Beyond quantifying the structure and connectivity of  
768 magma plumbing systems, seismic-based studies have shown that: (i) magma flow patterns mapped  
769 across entire sill-complexes indicate they can transport melt from source to surface over great  
770 lateral (>100's km) and vertical distances (10's km), potentially without significant input from  
771 dykes (Fig. 10a) (e.g., Thomson & Hutton, 2004; Cartwright & Hansen, 2006; Magee *et al.*, 2014;  
772 Magee *et al.*, 2016; Schofield *et al.*, 2017); and (ii) a variety of elastic and inelastic mechanisms can  
773 accommodate host rock deformation during magma emplacement, meaning that the location and  
774 size of ground deformation does not necessarily equal that of the forcing intrusion (e.g., Jackson *et*  
775 *al.*, 2013, Magee *et al.*, 2013b). Importantly, observations from seismic reflection data highlight that  
776 the lateral dimension should be considered when modelling the transit of magma in the crust, posing  
777 problems for the widely held and simple assumption that magma simply travels vertically from melt  
778 source to eruption site.

779 Seismic-based studies have also shown that direct comparison to active deformation  
780 structures can be informative. For example, through comparing mapped lava flows and structures

781 associated with the Alu dome to similar features observed in seismic reflection data (see section  
782 2.6), Magee *et al.*, (2017) concluded that the shallow-level sill likely has a saucer-shaped, as  
783 opposed to the sill-like tabular morphology inferred from an episode of deformation measured using  
784 InSAR (Figs 3c and d). Despite its benefits, it is important to remember that seismic reflection data  
785 typically reveal only the final geometry of the magma plumbing system. There thus remains a  
786 challenge in using these data to understand areas where deformation captures potentially transient,  
787 active processes, rather than structures resulting from (multiple) periods of intrusion and cooling  
788 (Reeves *et al.*, 2018). One potential and exciting way forward is the development of Virtual  
789 Reflection Seismic Profiling, by which microseismicity at active volcanoes may be used to image  
790 magma reservoirs and subsurface structure in 4D (Kim *et al.*, 2017). Although challenges exist in  
791 dataset integration, the imaging power afforded by modern seismic reflection data thus presents a  
792 unique opportunity to further unite field-, petrological-, geochemical-, and other geophysical-based  
793 analyses within more realistic structural frameworks (e.g., Figs 3, 11a and b). In our view, however,  
794 seismic reflection data are under-utilized in igneous research, remaining an unfamiliar technique to  
795 many Earth Scientists in the volcanic and magmatic community.

796

## 797 **2.7. Rock magnetism**

### 798 ***Technique***

799 Whilst seismic reflection data provide unique 3D images of ancient magma plumbing systems,  
800 which can be used to infer magma flow patterns across entire intrusion networks, we commonly  
801 lack sufficient data (e.g., boreholes) to test seismic-based hypotheses. It is therefore critical to  
802 compare seismic interpretations to field analogues where magma flow patterns, emplacement  
803 mechanics, and intrusion evolution can be investigated via other techniques. In this section, we  
804 examine how rock magnetic analyses can be used to systematically study magnetic mineralogy and  
805 petrofabrics, thereby illuminating the structure and history of igneous intrusions.



806           There are two principal types of rock magnetic study; magnetic remanence and magnetic  
807 susceptibility, where the total magnetisation ( $M$ ) of a rock is the sum of the magnetic remanence  
808 ( $M_{\text{rem}}$ ) and the induced magnetisation ( $M_{\text{ind}}$ ), which is a product of the susceptibility ( $K$ ) and  
809 applied field strength ( $H$ ) (Dunlop & Özdemir, 2001). Remanence carries a geological record of the  
810 various magnetisations acquired over time and is central to palaeomagnetic studies. However, we  
811 focus on magnetic fabric analysis, which relies on measurements of the anisotropy of magnetic  
812 susceptibility (AMS). The AMS signal of a rock carries information from all constituent grains.  
813 Although mineral phases that have a paramagnetic behaviour (i.e. they are weakly attracted to  
814 externally applied magnetic fields) volumetrically dominate most igneous rocks (e.g., olivine,  
815 clinopyroxene, biotite), ferromagnetic mineral phases (e.g., titanomagnetite) are highly susceptible  
816 to magnetization and therefore tend to dominate  $K$  (e.g., Dunlop & Özdemir, 2001; Biedermann *et*  
817 *al.*, 2014). Magnetic fabrics therefore typically reflect the preferential orientation of  
818 crystallographic axes (i.e. crystalline anisotropy), the shape-preferred orientation of individual  
819 crystals (i.e. shape anisotropy), and/or the alignment of closely spaced crystals (i.e. distribution  
820 anisotropy) belonging to Fe-bearing silicate and oxide phases (e.g., Voight & Kinoshita, 1907;  
821 Graham, 1954; Hrouda, 1982; Tarling & Hrouda, 1993; Dunlop & Özdemir, 2001). The principal  
822 axes of the magnetic fabrics measured by AMS can thus be related to the orientation, shape, and  
823 distribution of individual grains (i.e. the petrofabric) (e.g., Fig. 12a).

824           Regardless of whether mineral phases crystallise early or late, whereby their orientation and  
825 distribution typically mimics the earlier silicate framework, it is expected that the initial petrofabric  
826 developed in intrusive rocks will likely be sensitive to alignment of crystals during primary magma  
827 flow. However, it is also critical to recognise that later magmatic processes (e.g., convection and  
828 melt extraction) and syn- or post-emplacement tectonic deformation can modify or overprint  
829 primary magma flow fabrics during intrusion, solidification (i.e. mush development), or sub-solidus  
830 conditions (e.g., Borradaile & Henry, 1997; Bouchez, 1997; O'Driscoll *et al.*, 2015; Kavanagh *et*  
831 *al.*, 2018). Whilst anisotropy of magnetic susceptibility (AMS) can thus rapidly and accurately

832 detect weak or subtle mineral alignments within igneous intrusions, which may be attributable to  
833 magmatic and/or tectonic processes, evaluating the origin and evolution of petrofabric development  
834 requires additional information (e.g., Borradaile & Henry, 1997; Bouchez, 1997). For example,  
835 shape-preferred orientation analyses and comparison to visible flow indicators (e.g., intrusive steps  
836 and bridge structures) allow magma flow axes and directions that have been inferred from magnetic  
837 fabrics to be verified (e.g., Launeau & Cruden, 1998; Callot *et al.*, 2001; Magee *et al.*, 2012a). For a  
838 useful précis of AMS-related magnetic theory in igneous rocks, the reader is referred to early works  
839 by Balsey & Buddington (1960) and Khan (1962), and more recent summaries provided by Martín-  
840 Hernández *et al.*, (2004), O'Driscoll *et al.*, (2008), and O'Driscoll *et al.*, (2015).

841         The principle behind AMS relies on the measurement of the bulk susceptibility ( $K_m$ ) of a  
842 single sample in different orientations to determine the susceptibility anisotropy tensor, which  
843 relates the induced magnetisation ( $M_{ind}$ ) to the applied field ( $H$ ) in three dimensions (Tarling &  
844 Hrouda, 1993). The orientation and magnitude of the eigenvectors and eigenvalues of this tensor  
845 define an ellipsoid with three principal axes; the long axis of the ellipsoid,  $K_1$ , defines the magnetic  
846 lineation and the short axis,  $K_3$ , defines the normal (i.e. the pole) to the magnetic foliation plane  
847 ( $K_1$ – $K_2$ ; Fig. 12a) (Stacy *et al.*, 1960; Khan, 1962; Tarling & Hrouda, 1993). In order to interpret  
848 magnetic fabrics, it is important to determine the mineralogy of the phases carrying the magnetic  
849 signal because the composition, grainsize, and distribution of magnetically dominant minerals (e.g.,  
850 titanomagnetite) can control fabric orientation (e.g., Hargreaves *et al.*, 1991; Stephenson, 1994;  
851 Dunlop & Özdemir, 2001). In addition to primary crystallographic and textural controls on  
852 magnetic fabrics, subsequent oxidation of remaining melt and secondary hydrothermal alteration  
853 can affect the magnetic mineralogy and, thereby, the AMS signal (e.g., Trindade *et al.*, 2001;  
854 Stevenson *et al.*, 2007a). A variety of rock magnetic experiments are thus required to determine the  
855 magnetic mineralogy. The most widely used method involves measuring susceptibility, and thereby  
856 behaviour of magnetic materials, at varying temperatures ranging from  $-200^{\circ}\text{C}$  to  $700^{\circ}\text{C}$  (i.e.  
857 thermomagnetic analysis *sensu* Orlický, 1990; Hrouda *et al.*, 1997). For example, paramagnetic

858 materials (e.g., biotite) follow the Curie-Weiss law, whereby their susceptibility drops  
859 hyperbolically with increasing temperature. In contrast, the thermomagnetic curve of ferromagnetic  
860 materials (e.g., titanomagnetite) displays little change in susceptibility with temperature, apart from  
861 when characteristic crystallographic transitions occur (e.g., the Curie point for pure magnetite at  
862 ~580°C, Petrovský & Kapička, 2006) temperature. To determine the grainsize of ferromagnetic  
863 fraction in the magnetic susceptibility signal, the hysteretic property of the magnetisation is  
864 important (Dunlop, 2002). Other rock magnetic experiments (e.g., anisotropy of anhysteretic  
865 remanent magnetism (AARM) can be conducted to further isolate the relative importance of  
866 different paramagnetic and ferromagnetic phases (e.g., McCabe *et al.*, 1985; Richter & van der  
867 Pluijm, 1994; Kelso *et al.*, 2002).

868

## 869 ***Observations***

870 Having established the magnetic mineralogy, AMS fabrics can be interpreted. Even in weakly  
871 anisotropic igneous rocks (i.e. visually isotropic), particularly sheet intrusions, it is now accepted  
872 that the magnetic lineation and foliation can provide information on magma migration (e.g., flow  
873 direction) or regional and local strain (e.g., Hrouda, 1982; Knight & Walker, 1988; Rochette *et al.*,  
874 1992; Bouchez, 1997; Tauxe *et al.*, 1998; Callot *et al.*, 2001; Féménias *et al.*, 2004; Magee *et al.*,  
875 2012a). For example, comparisons to other indicators of magma flow (e.g., intrusive steps and  
876 visible mineral alignments) in sheet intrusions have shown that magnetic lineations commonly  
877 parallel the magma flow (e.g., Knight & Walker, 1988; Cruden & Launeau, 1994; Callot *et al.*,  
878 2001; Magee *et al.*, 2012a), whilst imbrication of elongate crystals induced by simple shear at  
879 intrusion margins define the sense of magma flow (Fig. 12b) (e.g., Knight & Walker, 1988;  
880 Hargraves *et al.*, 1991; Stephenson, 1994; Geoffroy *et al.*, 2002; Féménias *et al.*, 2004).  
881 Alternatively, contact-parallel magnetic fabrics generated during the formation and inflation of  
882 magma lobes can be used to determine flow and emplacement dynamics, even if other evidence for  
883 the presence of magma lobes is lacking (e.g., Fig. 12c) (Cruden *et al.*, 1999; Stevenson *et al.*,

2007a; Magee *et al.*, 2012b). Identifying changes in fabric orientation within or between individual sheet intrusions is also important because these variations suggest that deformation, imparted by either the emplacement of adjacent magma bodies or tectonic processes, did not significantly modify magma emplacement fabrics (e.g., Clemente *et al.*, 2007).

Post solidification textural modification and the possibility of overlap in tectonic and magmatic strain fields during protracted emplacement is a particular complication when studying granitoid and gabbroic plutons (e.g., Mamtani *et al.*, 2013; O'Driscoll *et al.*, 2015; Cheadle *et al.*, 2017). In fact, most early studies of granitoid emplacement using AMS, in conjunction with many other structural analysis tools, concluded that tectonic strain was the main source of subtle fabrics (e.g., Brun *et al.*, 1990; Bouchez, 1997; de Saint-Blanquat & Tikoff 1997; Neves *et al.*, 2003; Mamtani *et al.*, 2005). Although primary magma flow fabrics in granitic and gabbroic plutons may thus be overprinted, the magnetic fabrics characterised by AMS can still provide fundamental insights into emplacement mechanics (e.g., Stevenson *et al.*, 2007a; Petronis *et al.*, 2012) and magma/mush evolution (e.g., formation of layering; O'Driscoll *et al.*, 2015).

### ***Implications and integration***

Overall, AMS has provided vital magma flow and evolution information that has helped to understand mafic and silicic magma plumbing systems (e.g., Knight & Walker, 1988; Ernst & Baragar, 1992; Glen *et al.*, 1997; Aubourg *et al.*, 2008; Petronis *et al.*, 2013; Petronis *et al.*, 2015). Critical insights emanating from these AMS studies have revealed that: (i) flow trajectories predicted by classic emplacement models (e.g., for ring dykes and cone sheets) are not always consistent with measured AMS fabrics and supporting data, which thereby call into question the application of such models (e.g., Stevenson *et al.*, 2007b; Magee *et al.*, 2012a); (ii) lateral magma flow is recorded in many shallow, planar intrusions associated with volcanic magma plumbing systems (e.g., Ernst & Baragar, 1992; Cruden & Laneau, 1994; Cruden *et al.*, 1999; Herrero-Bervera *et al.*, 2001; Magee *et al.*, 2012a; Petronis *et al.*, 2013; Petronis *et al.*, 2015); and (iii)

910 plutons, particularly those with a granitic composition, commonly consist of incrementally  
911 emplaced magma pulses that often develop lobate geometries (e.g., Fig. 12c) (e.g., Stevenson et al.,  
912 2007a). Analysing AMS fabrics from layered mafic-ultramafic intrusions can also provide evidence  
913 for magma reservoir processes, including crystal settling, or post-cumulus modification of crystal  
914 mushes (O'Driscoll *et al.*, 2008; O'Driscoll *et al.*, 2015). Importantly, AMS and related analyses  
915 provide robust, testable, and repeatable methods to constrain subtle shape and crystallographic  
916 orientations of crystals in igneous rocks. Rock magnetic instrumentation technology continues to  
917 advance with better automation of measurement protocols, sensitivity of measurements, and a  
918 greater ability to unravel contributors to the AMS signal. The direction and scope of these  
919 developments are improving the holistic integration of AMS with other structural, microstructural,  
920 geophysical, petrological and geochemical techniques, promising to advance our understanding of  
921 magmatism and crustal evolution.

922

### 923 **3. Future advances**

924 Our understanding of magma plumbing system structure and evolution has been significantly  
925 enhanced by the geophysical techniques described above. We have demonstrated that there is scope  
926 for advancement within individual methodologies and through the integration of different  
927 techniques, particularly involving the synthesis of geophysical, petrological, and geochemical data.  
928 In this section, we discuss two new techniques that will potentially revolutionize our understanding  
929 of magma plumbing systems. We also briefly discuss how integration of geophysical data with  
930 numerical modelling can enhance our knowledge of reservoir construction and evolution.

931

#### 932 **3.1. Full-Waveform Inversion**

##### 933 ***Technique***

934 We have demonstrated that seismic reflection data can provide unique insight into the 3D structure  
935 of magma plumbing systems (e.g., see review by Magee *et al.*, 2016). In addition to using seismic

936 reflection data to image the subsurface, we can also invert the measured travel-times of reflected  
937 acoustic energy to model subsurface P-wave velocities. Full-waveform inversion (FWI) is a rapidly  
938 developing technology using active source seismic data to generate models that reproduce both the  
939 travel-times and full waveform of the arriving wavefield, thereby matching observed seismic data  
940 (Tarantola, 1984). Because FWI considers the full wavefield, as opposed to conventional techniques  
941 that only model travel-times, it is a technique capable of recovering high-resolution models of  
942 subsurface P-wave velocities and other physical properties (Warner *et al.*, 2013; Routh *et al.*, 2017).  
943 The FWI technique begins with a best-guess starting velocity model for the subsurface geology,  
944 which is then iteratively updated using a local linearized inversion until the observed seismic data is  
945 matched (Virieux & Operto, 2009). FWI is much more computationally expensive than travel-time  
946 tomography, as a full-physics implementation of the wave equation is required to generate the  
947 predicted seismic data at all energy source and receiver locations for each iteration (Routh *et al.*,  
948 2017). FWI, however, has the advantage of being able to resolve much finer-scale structure than  
949 conventional techniques.

950

## 951 ***Observations***

952 To date, 3D FWI has principally been applied within the petroleum sector to obtain high-resolution  
953 velocity models that can be used to improve depth-migrated (i.e. travel-time is converted to depth in  
954 metres) reflection images of petroleum reservoirs and their overburden (Sirgue *et al.*, 2010; Vigh *et al.*,  
955 2010; Warner *et al.*, 2013; Kapoor *et al.*, 2013; Routh *et al.*, 2017). FWI can also produce  
956 interpretable, quantitative models of physical properties of rocks in the subsurface that can be  
957 related directly to compaction, permeability, and overpressure as measured in subsurface boreholes  
958 (Lazaratos *et al.*, 2011; Mancini *et al.*, 2015). Of relevance here is that mafic intrusions, which  
959 appear as high-amplitude reflections in seismic reflection data (Figs 10 and 11a), are recovered as  
960 high-velocity features in FWI velocity models (Fig. 13) (Mancini *et al.*, 2015; Kalincheva *et al.*,  
961 2017). For example, successful application of 3D FWI to a marine ocean bottom seismometer

962 dataset acquired across the Endeavour segment of the Juan de Fuca Ridge led to generation of a  
963 velocity model that had a resolution up to four times greater than travel-time tomography (Morgan  
964 *et al.*, 2016). Within this new, high-resolution velocity model, several velocity anomalies were  
965 identified and interpreted to indicate localized magma recharge of the axial reservoir, induced  
966 seismogenic cracking, and increased permeability (Arnoux *et al.*, 2017).

967

### 968 ***Implications and integration***

969 Active magma plumbing systems comprise a complex network of interconnected conduits and  
970 reservoirs with variable geometries and sizes, which likely contain magmatic vapour-rich, liquid-  
971 rich, and mush-zones (Christopher *et al.*, 2015). These intrusions will all be associated with reduced  
972 P-wave velocities, which could be resolved in high-resolution, 3D FWI datasets as supported by  
973 successes in the fine-scale imaging of: (i) low-velocity gas clouds (Warner *et al.*, 2013); (ii) axial  
974 reservoirs at an oceanic spreading centre (Arnoux *et al.*, 2017); (iii) relatively narrow, low-velocity  
975 fault zones within an antiform (Morgan *et al.*, 2013); and (iv) a subduction zone using 2D FWI  
976 (Kamei *et al.*, 2012). A suite of synthetic tests has been performed to investigate whether 3D FWI  
977 could be applied to better understand magma plumbing systems (Morgan *et al.*, 2013). These tests  
978 indicate that it is possible to recover high-resolution models of P-wave velocity beneath volcanoes,  
979 which can then be used to better determine where magma/mush is stored beneath the surface. In  
980 particular, these synthetic tests suggest that FWI could be used to: (i) distinguish between  
981 continuous zones of mush and individual magma reservoirs; (ii) image sills and conduits of magma  
982 and/or fluids that are a few 10s metres across (Fig. 13); and (iii) image the deeper (lower-crustal)  
983 part of the magma system. We therefore consider that 3D FWI affords an unprecedented  
984 opportunity to obtain high-resolution images of actual magma plumbing systems beneath active  
985 volcanoes. To this end, the ongoing PROTEUS (Plumbing Reservoirs Of The Earth Under  
986 Santorini) experiment was specifically designed to use 3D FWI to investigate the Santorini magma  
987 plumbing system (Hooft *et al.*, 2017).

988

## 989 **3.2. Unmanned Aerial Vehicle photogrammetry**

### 990 *Technique*

991 Despite major advances in satellite-based remote sensing systems and aeromagnetic surveys, very  
992 high-resolution (i.e., mm–cm scale ground sampling distance) imagery of dykes and other igneous  
993 intrusions has been limited to low altitude aerial photography. This in turn has created a critical  
994 scale gap in intrusion studies, which range from <1 mm at thin section scale to the metres to 100's  
995 of metres scale provided by outcrop analysis, conventional remote sensing, and geophysical data.  
996 Fortunately, the emerging capability of unmanned aerial vehicle (UAV) photogrammetry fills this  
997 gap (e.g., Eisenbeiss, 2009; Westoby *et al.*, 2012; Bemis *et al.*, 2015; Eide *et al.*, 2017b). It is also  
998 noteworthy that several studies have demonstrated that digital photogrammetry can deliver high  
999 quality datasets with accuracies similar to more established laser scanning techniques (e.g., Leberl  
1000 *et al.*, 2010; Hodgetts, 2013; Thiele *et al.*, 2015).

1001 The basic setup required to carry out UAV (or drone) photogrammetry is commercially  
1002 available and relatively inexpensive, comprising a fixed wing or rotary wing UAV, a digital camera,  
1003 and access to a suitable digital photogrammetry software package (e.g., Agisoft Photoscan Pro,  
1004 Pix4Dmapper Pro, VisualSFM). UAV photogrammetry combines a simple and cost-effective  
1005 method to acquire geospatially referenced, overlapping digital aerial images, from which structure-  
1006 from-motion algorithms can generate spatial 3D datasets (Bemis *et al.*, 2014; Vollgger & Cruden,  
1007 2016). Such an approach can be used for high spatial resolution mapping of all types of well-  
1008 exposed igneous intrusions. The resulting data greatly enhance the effectiveness of traditional field  
1009 mapping, particularly the characterisation of contact relationships and internal and external structure  
1010 (e.g., fractures, fabrics, and phase distributions) of intrusive rocks, complementing AMS and  
1011 petrological analyses.

1012

### 1013 *Observations*



1014 Aa photogrammetric workflow was applied to examine a swarm of 5 cm to 1 m wide Palaeogene  
1015 dolerite and dacite dykes exposed on coastal outcrops at Bingie Bingie Point, SE Australia (Fig.  
1016 14). The orthophotograph of the entire wave-cut platform shows the distribution of the Palaeogene  
1017 dolerite and dacite dykes and their Devonian host rock lithologies, including a prominent  
1018 moderately NE-dipping aplite dyke (Fig. 14a). Linear ENE-WSW linear terrain features pick out  
1019 the traces of dyke-parallel joints (Fig. 14a). The Palaeogene dykes trend  $063^{\circ}$  parallel to a major set  
1020 of joints in the country rock that likely formed contemporaneously with syn-dyking extension (Fig.  
1021 14b). Subsidiary joint sets trend NNW-SSE, sub-perpendicular to the Palaeogene dykes, N-S and E-  
1022 W (Fig. 14b). The Palaeogene dykes display considerable structural complexity such as bridge  
1023 structures, intrusive steps and apophyses (Fig. 14c). Where present, the steps mostly occur where  
1024 dykes cross country rock contacts (e.g., the aplite-tonalite contact in the NE; Fig. 14c).

1025

### 1026 ***Implications and integration***

1027 Data such as the orthophotograph collected at Bingie Bingie Point indicate that high-resolution  
1028 structural and lithological mapping and measurement can be carried out much more rapidly than by  
1029 traditional survey methods (e.g., plane table or grid mapping). However, the use of conventional  
1030 RGB cameras restricts the resulting image data to reflected visible light. Future applications will  
1031 include the deployment of multispectral and hyperspectral sensors (infrared to short wave infrared  
1032 to thermal infrared) as well as potential field geophysical or geodetic instruments (e.g., Sparks,  
1033 2012). A further challenge for UAV applications in many countries concerns the regulatory  
1034 framework around the use of drones for research. The global trend is moving to require non-  
1035 recreational UAV operators to have remotely piloted aircraft licences and for the associated  
1036 organisation to be certified for UAV operations. Innovations in sensor types and design, attachment  
1037 of geophysical instruments, machine learning, and integration with complementary techniques such  
1038 as AMS will open up new avenues for UAV applications in the study of magma plumbing systems.

1039

### 1040 **3.4. Numerical modelling of magma reservoir processes constrained by geophysical data**

1041 Geophysical imaging of both active and ancient magma plumbing systems is delivering new  
1042 insights into the 3D geometry of reservoirs, the timing and rates of melt and magma transport, the  
1043 pathways followed by magmas as they ascend through the crust, and typical stored melt fractions in  
1044 mushes. These data can be used to constrain and calibrate numerical models of reservoir processes.  
1045 Numerical models are used ubiquitously to understand and predict the behaviour of other  
1046 subsurface crustal reservoirs, such as hydrocarbon reservoirs, groundwater resources, and targets for  
1047 geological CO<sub>2</sub> storage (e.g., Chen *et al.*, 2003; Class *et al.*, 2009; Dean & Chen, 2011). However,  
1048 there has been relatively little focus to date on developing numerical models for magma/mush  
1049 reservoirs. Yet such models can integrate across different data sources and types, provide  
1050 quantitative estimates of rates, volumes and timescales, and provide a framework for data  
1051 interpretation. For example, numerical modelling of heat transfer within the plumbing system at  
1052 Okmok Volcano in Alaska, which was informed by analytical models of geodetic data and  
1053 estimated magma compositions of erupted material, allowed estimation of the role magma injection,  
1054 crystallisation, and degassing processes had on volume changes over time (Caricchi *et al.*, 2014).  
1055 Numerical thermal modelling has also helped interpret seismic data from the Soufrière Hills  
1056 Volcano, Montserrat, suggesting higher melt fraction in the underlying magma reservoir than was  
1057 inferred from seismic data alone (Paulatto *et al.*, 2012). More recent numerical models focus on  
1058 crystal mushes, evaluating melt transport and reaction at low melt fractions, and these show that  
1059 temperature and melt fraction in mushes can be decoupled; i.e. maximum temperature occurs close  
1060 to the centre of the reservoir but maximum melt fraction occurs close to the top (Solano *et al.*,  
1061 2014). This decoupling impacts how seismic velocities and electrical conductivities will be  
1062 modified within the mush (Solano *et al.*, 2014). Other numerical models show the important role  
1063 played by exsolution, crystallisation, and the viscoelastic response of the crust in driving magma  
1064 mobilisation in and eruption from shallow reservoirs (e.g., Degruyter & Huber, 2014; Parmigiani *et*  
1065 *al.*, 2016), as well as providing insights into the mixing mechanisms of melt and crystals in mushes

1066 (Bergantz *et al.*, 2015). However, most models to date have a lower dimensionality (zero dimension  
1067 box models, or one/two dimensions) and capture only a small subset of the key physical and  
1068 chemical processes that are likely to occur in crustal magma reservoirs or crystal mushes.  
1069 Moreover, few studies have integrated modelling with geophysical data (cf. Gutierrez *et al.*, 2013).  
1070 This is in marked contrast to the 3D modelling routinely undertaken of other crustal reservoirs (e.g.,  
1071 hydrocarbon reservoirs), which is commonly integrated with and delimited by geophysical data.  
1072 There is thus significant scope for improved, and integrated, numerical modelling of crustal magma  
1073 reservoirs.

1074

#### 1075 **4. Conclusions**

1076 Determining the structure of magma plumbing systems is critical to understanding where melt and  
1077 magma is stored in the crust, which can influence the location of volcanic eruptions and economic  
1078 ore deposits, providing an important framework for interpreting the physical and chemical evolution  
1079 of magma from petrological and geochemical datasets. Geophysical techniques have revealed  
1080 unique insights into the architecture of active and ancient magma plumbing systems, which when  
1081 integrated with traditional structural, petrological and geochemical results has yielded exciting  
1082 advances in our understanding of magmatic processes. However, divisions between communities  
1083 applying these methodologies still exist, contributing to diverging views on the nature of magma  
1084 plumbing systems. To help promote collaboration, we have reviewed a range of geophysical  
1085 techniques and discussed how they could be integrated with structural, petrological and  
1086 geochemical datasets to answer outstanding questions in the volcanological community. In  
1087 particular, we demonstrate how a range geophysical techniques can be applied to track melt  
1088 migration in near real-time, map entire intrusion networks in 3D, examine magma emplacement  
1089 mechanics, and understand the evolution of crystal mushes. For example, Interferometric Synthetic  
1090 Aperture Radar (InSAR) allows measurement of the development of active magmatic systems by  
1091 successive intrusion, the vertical and lateral movements of magma, and the relationship between

1092 magma plumbing system dynamics and eruption. Seismicity beneath volcanoes can, when the  
1093 magma interacts dynamically with the host rock, illuminate in high-resolution the time and spatial  
1094 scales of the motion of magma and hydrothermal fluids. Seismic imaging of magma plumbing  
1095 systems allows the spatial distribution of melt and magma to be determined whilst the inclusion of  
1096 anisotropy within seismic techniques even allows sub-seismic wavelength features to be identified.  
1097 Gravimetry can characterise the distribution and redistribution of mass (e.g., magma) in the  
1098 subsurface over high spatial and temporal resolutions, helping to reveal the structure and  
1099 composition of magma plumbing systems and the source(s) of volcano deformation.  
1100 Electromagnetic methods, particularly magnetotellurics, can identify fluids within magmatic  
1101 systems (e.g., melt, magma, and hydrothermal fluids). Seismic reflection data provide  
1102 unprecedented 3D images of ancient magma plumbing systems and has revealed that laterally  
1103 extensive, interconnected networks of sills and inclined sheets can play a pivotal role in transporting  
1104 magma through the crust to eruption sites potentially located >100 km away from the melt source.  
1105 Rock magnetics can provide fabric data pertaining to magma flow, deformation or crystallisation.  
1106 All these methodologies discussed have provided unique insights into the structure of igneous  
1107 intrusions and, through integration with petrological and geochemical datasets, are beginning to  
1108 help unravel the entire evolution of magma plumbing systems. In addition to the ongoing  
1109 application and advancement of these geophysical techniques, emerging methodologies look set to  
1110 radically improve our understanding of magma plumbing systems. For example, full-waveform  
1111 inversion can image and characterise physical properties across plumbing systems at an  
1112 unprecedented resolution, whereas unmanned aerial vehicle photogrammetry provides a tool for  
1113 high spatial resolution of outcrop scale intrusions that bridges the scale gap between seismic  
1114 reflection data and traditional mapping of magma plumbing systems. The geophysical techniques  
1115 discussed also provide critical constraints on input parameters for numerical modelling. Overall, we  
1116 consider that the future of magma plumbing system studies will benefit greatly from the synthesis  
1117 of geophysics and more traditional petrological and geochemical approaches.

1118

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1968

## 1969 **7. Figure captions**

1970

1971 Figure 1: Schematic of a vertically extensive, transcrustal magma plumbing system involving  
1972 transient, interconnected, relatively low-volume tabular magma intrusions (e.g., dykes, sills, and  
1973 laccoliths) within a crystal mush (based on Cashman *et al.*, 2017; Cruden *et al.*, 2018).

1974

1975 Figure 2: (A) Interferograms showing fringes caused by the pressurisation of a point source directly  
1976 beneath a stratovolcano from both ascending and descending satellite lines of sight. Note that the  
1977 centre of the fringes are slightly offset from the summit of the volcano (marked by a black triangle).  
1978 (B) Typical fringe patterns for analytical deformation sources in an elastic half space from  
1979 ascending satellite geometry: (i) Mogi source at 5 km depth; (ii) dyke extending between depths of  
1980 3 and 9 km; (iii) rectangular sill; and (iv) a penny-shaped horizontal crack both at 5 km depth.

1981

1982 Figure 3: (A) Ascending line of sight (LOS) co-eruptive interferogram from the 2008 basalt lava  
1983 extrusion between the Alu and Alu South domes and the Dalafilla stratovolcano (modified from  
1984 Pagli *et al.*, 2012). (B) Inversion of uplift and subsidence patterns, recorded by InSAR during the  
1985 2008 basalt lava eruption at the Alu dome in the Danakil Depression, suggested ground deformation  
1986 could be attributed to a combination of: (i) deflation of a reservoir, modelled as a Mogi source, at  
1987 ~4 km depth; (ii) inflation and deflation of a tabular sill at ~1 km depth; and (iii) opening of a dyke  
1988 beneath the eruptive fissure (Figs 3A and B) (Pagli *et al.*, 2012). See Figure 3A for location. (C)  
1989 Geological map showing that lava flows radiate out from Alu and originate from the periphery of  
1990 the dome, which is cross-cut by an array of randomly oriented faults (modified from Magee *et al.*,  
1991 2017). (D) Magee *et al.*, (2017) inferred Alu is underlain by a saucer-shaped sill plumbing system,  
1992 based on field observations and comparison to seismic reflection data, not a tabular sill (Fig. 3B).

1993

1994 Figure 4: Example of integrating seismology and petrology to constrain time-scales of magma  
1995 storage and recharge (from Saunders *et al.*, 2012). Calculated Fe-Mg diffusion time scales of  
1996 orthopyroxene crystals compared to monitoring data for the same eruptive period for Mount St.  
1997 Helens. (A) The seismic record of depth against time of the 1980–1986 eruption sequence. (B)  
1998 Measured flux of SO<sub>2</sub> gas. (C) Calculated age of orthopyroxene rim growth binned by month for the  
1999 entire population. The age recorded is the month in which the orthopyroxene rim growth was  
2000 triggered by magmatic perturbation. The black line displays the running average (over five points,

equivalent to the average calculated uncertainty in calculated time scales) of all the data. The peaks in the diffusion time series correspond to episodes of deep seismicity in 1980 and 1982 and to elevated SO<sub>2</sub> flux in 1980 and possibly 1982. (D) Running average of the orthopyroxene rim time scales, displaying reverse zonation (Mg-rich rims) in blue and normal zonation (Fe-rich rims) in green. There are reverse zonation peaks in the early 1980, probably due to rejuvenation of the magma system by hotter pulses, whereas Fe-rich rims are more dominant from 1982 on. Vertical dashed grey lines represent the volcanic eruptions.

Figure 5: Plot of melt inclusion saturation and earthquake hypocentre depths, which suggest magma storage occurred at 1–5 km depths, beneath the Dabbahu volcanic system in Afar, Ethiopia (modified from Field *et al.*, 2012). Melt inclusion data obtained from analyses of alkali feldspar, clinopyroxene, and olivine phenocrysts within Dabbahu lavas <8 Kyr (Field *et al.*, 2012). Earthquake data recorded during the 2005 dyke event (Ebinger *et al.*, 2008).

Figure 6: (A) P-wave ( $V_p$ ) tomography beneath Montserrat (black outline), highlighting the location of fast and slow seismic velocity anomalies (i.e. >6% faster or slower than average) relative to the location of the Silver Hills (SH), Central Hills (CH), and Soufrière Hills (SHV) volcanoes (modified from Shalev *et al.*, 2010). The fast velocity anomalies, interpreted to represent solidified andesitic intrusions underlie the volcanoes (Shalev *et al.*, 2010).

Figure 7: Static and dynamic gravimetric investigations of two active silicic magmatic systems in the Andes: Uturuncu volcano (Bolivia; A, C, and E) and the Laguna del Maule volcanic field (Chile; B, D, and F). (A) 3D view of the isosurface corresponding to the -120 kg m<sup>3</sup> density contrast beneath Uturuncu volcano, derived from Bouguer gravity data, interpreted to reflect a large (~750 km<sup>3</sup>) plumbing system composed of a lower (<10 km) partially molten reservoir and upper, fractured and fluid-bearing solidified intrusions above sea level (after del Potro *et al.*, 2013). (B) 3D



view of the  $-600 \text{ kg m}^3$  density contrast isosurface beneath the Laguna del Maule, which is interpreted to define a magma reservoir ( $>50 \%$  melt) within a larger region of a crystal mush system; the 2D planes show slices through the dataset (Miller *et al.*, 2017). Elevation above sea level (a.s.l.) shown. See Figure 7D for area of data coverage. (C) Map of the 55 km long, dynamic gravity network (white circles) installed to track changes in gravity over time and space at Uturuncu volcano between 2010 and 2013 (modified from Gottsmann *et al.*, 2017). (D) Spatio-temporal residual gravity changes at Laguna del Maule recorded from 2013–2014, after correcting for deformation effects (modified from Miller *et al.*, 2017). (E) Gravity and deformation data, recorded from Uturuncu from 2010–2013, plotted against the measured free-air gravity gradient (solid red line) and associated errors (broken red lines) (modified from Gottsmann *et al.*, 2017). The data follow the gradient and are indicative of a subsurface density change as a cause of the uplift, possibly reflecting the release of fluids from a large deep-seated magma reservoir (i.e. the Altiplano-Puna Magmatic Body; Chmielowski *et al.*, 1999) through the vertically extensive crystal mush system shown in (A) (Gottsmann *et al.*, 2017). (F) Plot of gravity against horizontal distance for the source centre at Laguna del Maule (modified from Miller *et al.*, 2017). The increase in gravity of up to  $120 \mu\text{Gal}$  is explained by a hydrothermal fluid injection focused along a fault system, shown in (D), at 1.5–2 km depth as a result of a deeper seated magma injection, and is best modelled by a vertical rectangular prism source.

Figure 8: Gravity changes and deformation at the restless Long Valley caldera. (A) Map of the Long Valley caldera, California, USA, which hosts a resurgent dome (black outline), to highlight changes in residual gravity between 1982 and 1999 (modified from Tizzani *et al.*, 2009). (B) Plot of ground uplift and residual gravity changes with radial distance from the centre of the resurgent dome in (A) (modified from Tizzani *et al.*, 2009). The correlation between uplift and positive gravity residuals across the resurgent dome indicates ground deformation was instigated by intrusion of magma (Tizzani *et al.*, 2009).

2053

2054 Figure 9: (A) Map showing MT stations deployed around Volcán Uturuncu (U) and Volcán  
2055 Quetena (Q), relative to areas of uplift and subsidence (modified from Comeau *et al.*, 2015). The  
2056 white box shows area of modelled 3D MT data (Comeau *et al.*, 2015). (B) Regional 2D  
2057 magnetotelluric line through the Altiplano-Puna magma body (APMB) highlighting the position of  
2058 Volcán Uturuncu (modified from Comeau *et al.*, 2015). The APMB corresponds to a large,  
2059 conductive (i.e. low-resistivity) body (Comeau *et al.*, 2015; Comeau *et al.*, 2016). Above the APMB  
2060 are other areas of low-resistivity (e.g., C4) that are likely upper crustal magma reservoirs and dykes  
2061 (Comeau *et al.*, 2016). C1–C7 and R1–R2 identify discrete zones of marked conductivity or  
2062 resistivity, respectively (see Comeau *et al.*, 2015; Comeau *et al.*, 2016 for details). The white box  
2063 shows area of modelled 3D MT data (Comeau *et al.*, 2015). See Figure 9A for location.

2064

2065 Figure 10: (A) Interpreted seismic section and geological map showing the distribution of and  
2066 connectivity between sills within the Faroe-Shetland Basin (modified from Schofield *et al.*, 2017).  
2067 Mapping of magma flow patterns within individual sills reveals that the sill-complex facilitates  
2068 extensive vertical and lateral magma transport. Magma was fed into the sedimentary basin via  
2069 basement-involved faults. TWT = two-way travel time. (B) Interpreted seismic section and  
2070 geological map describing the spatial relationship between volcanoes/vents and sills, inferred to  
2071 represent the magma plumbing system, emplaced at ~42 Ma (modified from Jackson *et al.*, 2013;  
2072 Magee *et al.*, 2013a). Sills are laterally offset from the volcanoes/vents summits. No ‘magma  
2073 chambers’ are observed in the seismic data, which images down to ~8 s TWT (i.e. ~>10 km)  
2074 (Magee *et al.*, 2013a).

2075

2076 Figure 11: (A) Interpreted seismic section from the Exmouth Sub-basin offshore NW Australia,  
2077 which images a saucer-shaped sill that is overlain by a forced fold and feeds a small vent from its  
2078 inclined limb (modified from Magee *et al.*, 2013b). See Figure 11B for line location). (B) Time-

2079 structure map of the folded horizon (thick black line) in (A), highlighting fault traces and vent  
2080 locations and thicknesses (modified from Magee *et al.*, 2013b). (C) Seismic section from the  
2081 Farsund Basin, offshore southern Norway, which images part of a dyke-swarm that has been rotated  
2082 by basin flexure post-emplacement (modified from Phillips *et al.*, 2017).

2083  
2084 Figure 12: (A) At the sample scale, all magnetic grains create a magnetic fabric. (i) Dominantly  
2085 prolate fabric, where  $K_2$  and  $K_3$  are least certain and form a girdle. Only the magnetic lineation ( $K_1$ )  
2086 can be confidently determined. (ii) When  $K_1 > K_2 > K_3$ , both a foliation ( $K_1$ – $K_2$ ) and a lineation ( $K_1$ )  
2087 may be discerned, defining a triaxial fabric. (iii) When  $K_1$  and  $K_2$  are equally uncertain and form a  
2088 girdle,  $K_3$  is perpendicular to a foliation. (B) Schematic representation of how magma flow within a  
2089 planar sheet intrusion can produce imbricated magnetic fabrics at its margins, the closure of which  
2090 define the magma flow direction (after Féménias *et al.*, 2004). (C) AMS data and interpretations  
2091 from part of the Trawenagh Bay Granite, NW Ireland (adapted from Stevenson *et al.*, 2007a). (i)  
2092 AMS foliation traces are shown in blue and lineation traces in red. Lobes were defined in this  
2093 intrusion based on foliations curving around a lineation axis. In some lobes, the magnetic lineation  
2094 trend was parallel to this axis, whilst in others they tended to splay or converge down flow. (ii) 3D  
2095 sketch showing the geometry of three of the lobes (numbered in part i).

2096  
2097 Figure 13: (A) Starting model derived from smoothed, pre-stack, time-migrated (PSTM) stacking  
2098 velocities. (B) Final 2D FWI-derived velocity model obtained using 10 km streamer data and  
2099 inversion frequencies of between 2.5 and 24 Hz. (C) FWI velocity model overlain by the 2D pre-  
2100 stack, depth-migrated (PSDM) section. Strong irregular reflections in the lower half of the section  
2101 are from basaltic intrusions, which appear as high-velocity anomalies in the FWI velocity model.  
2102 Both the FWI velocity model and the PSDM pick out a major unconformity, and show shallow  
2103 channels in the upper parts of the section (redrawn from Kalincheva *et al.*, 2017).

2105 Figure 14: (A) UAV orthophotograph of the wave cut platform at Bingie Point, NSW, Australia  
2106 showing the distribution of Palaeogene dolerite (Dol) and dacite (Dac) dykes within Devonian  
2107 tonalite (Ton), diorite (Di), and aplite (Ap) host rocks. (B) Circular histogram of joint sets measured  
2108 in the Devonian rocks from the orthophotograph; the dominant (purple) set is parallel to and likely  
2109 contemporaneous with the Palaeogene dykes. (C) Annotated close-up image highlighting dykes and  
2110 structural features. The northern dacite dyke shows two broken bridge (BB) structures, whilst the  
2111 central dolerite dyke displays prominent step structures (S). Narrow apophyses are also associated  
2112 with the broken bridges and steps.

Figure 1

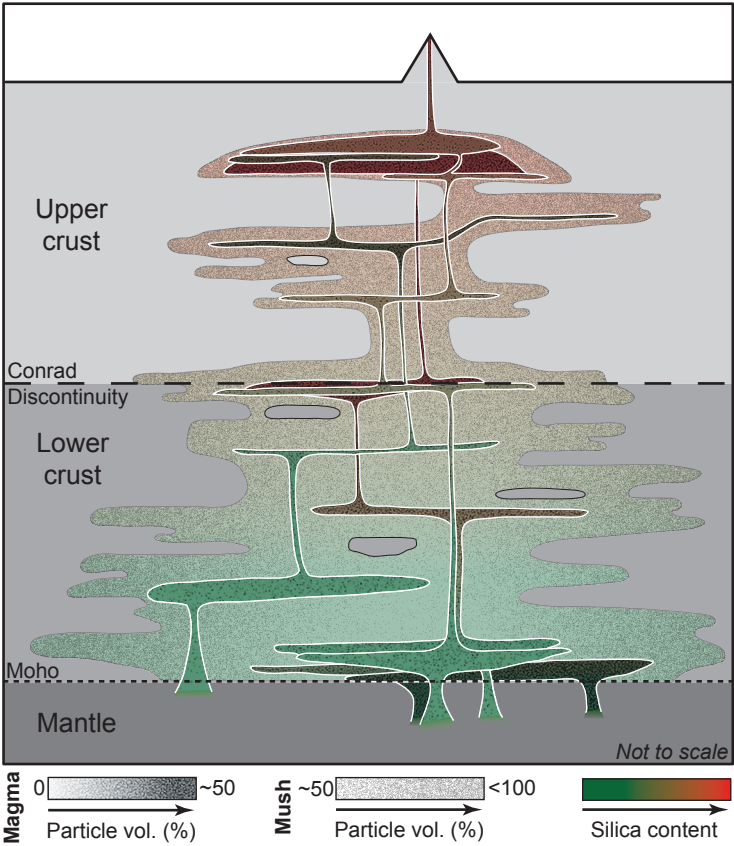
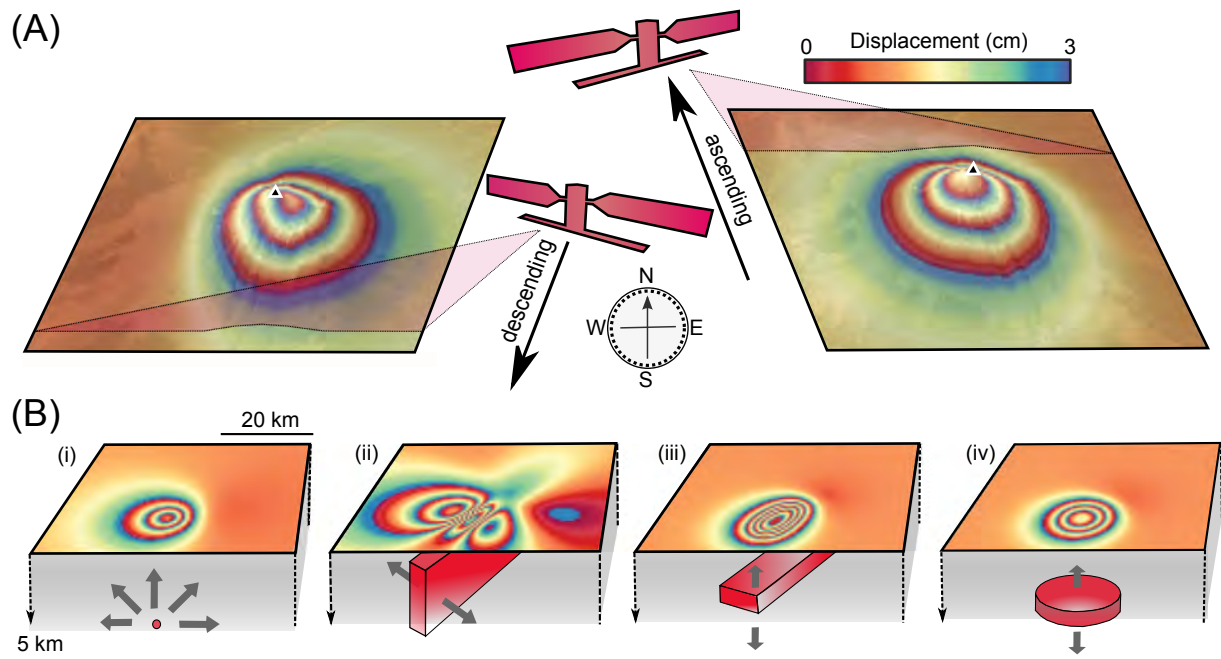


Figure 2



**Figure 3**

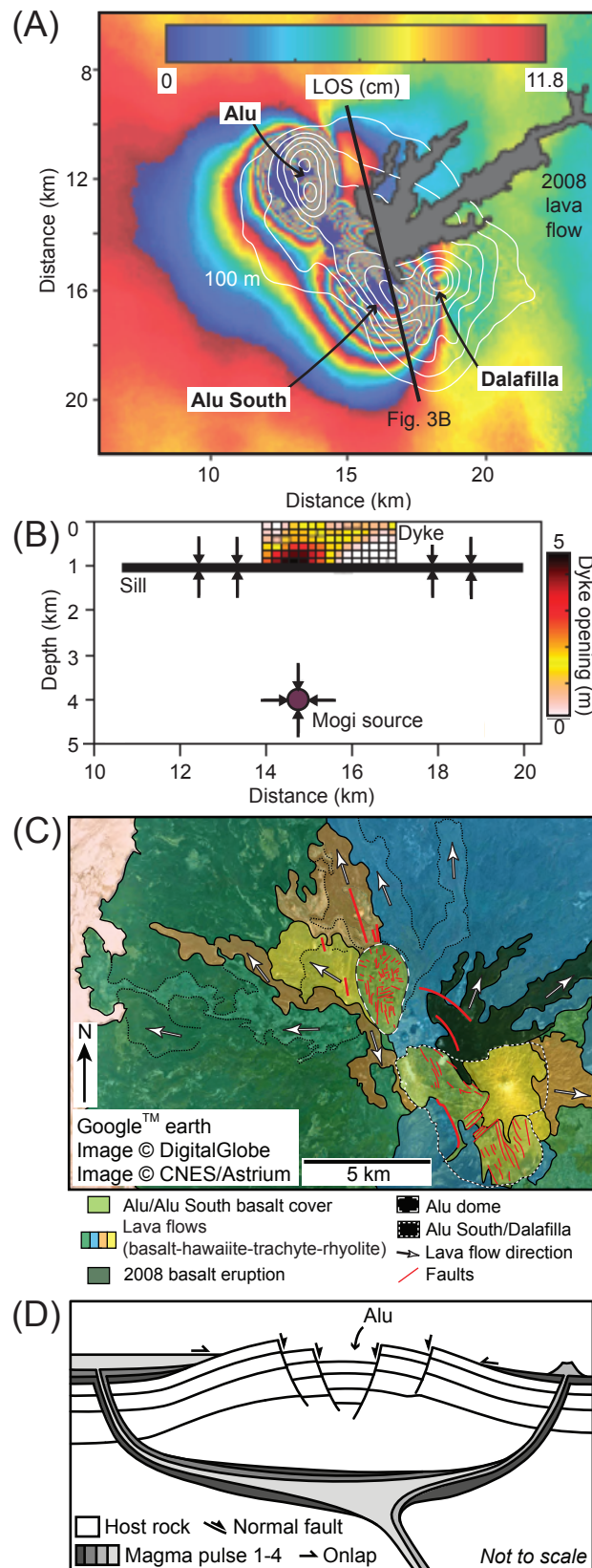


Figure 4

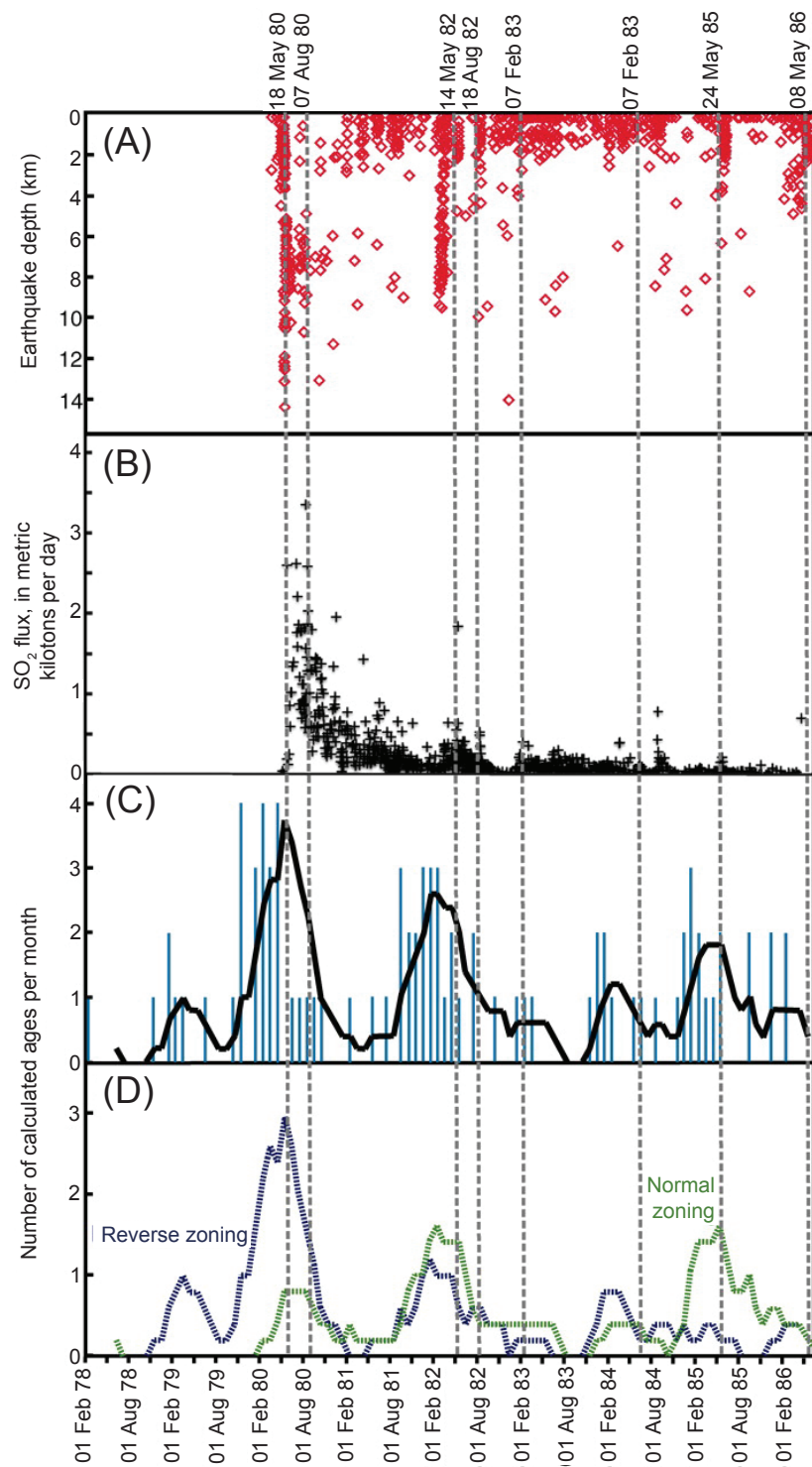




Figure 5

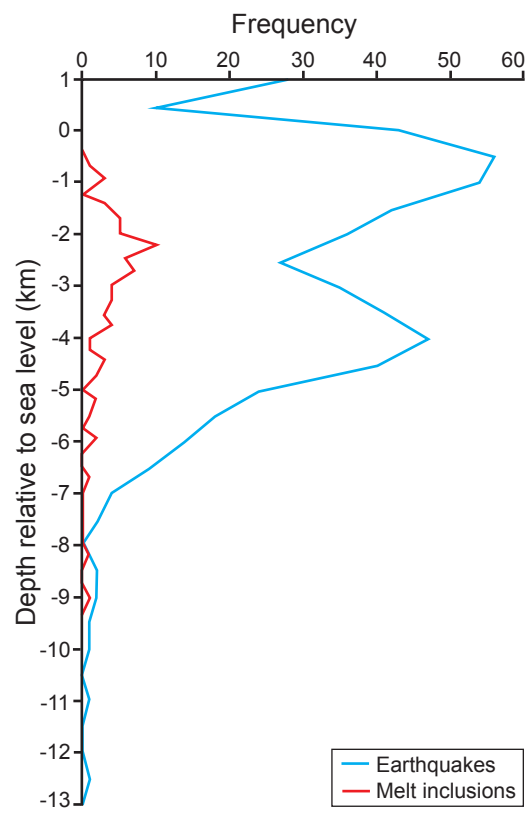


Figure 6

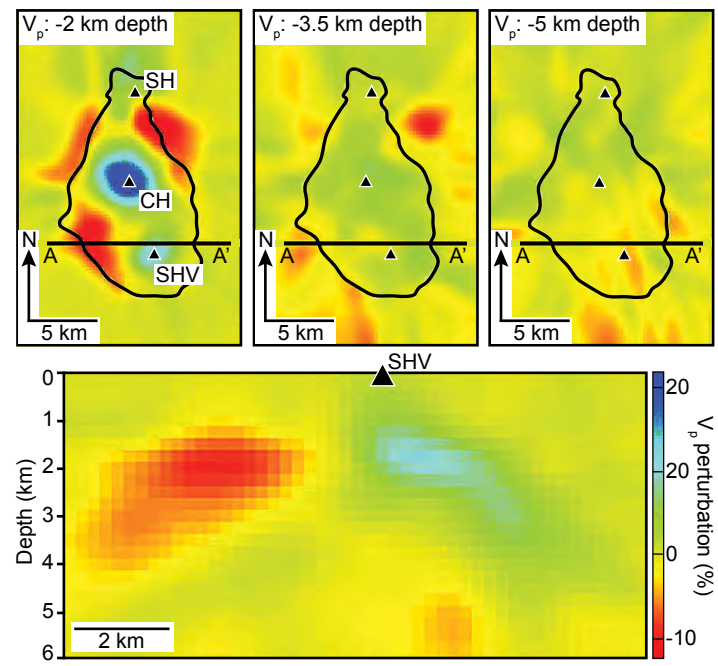
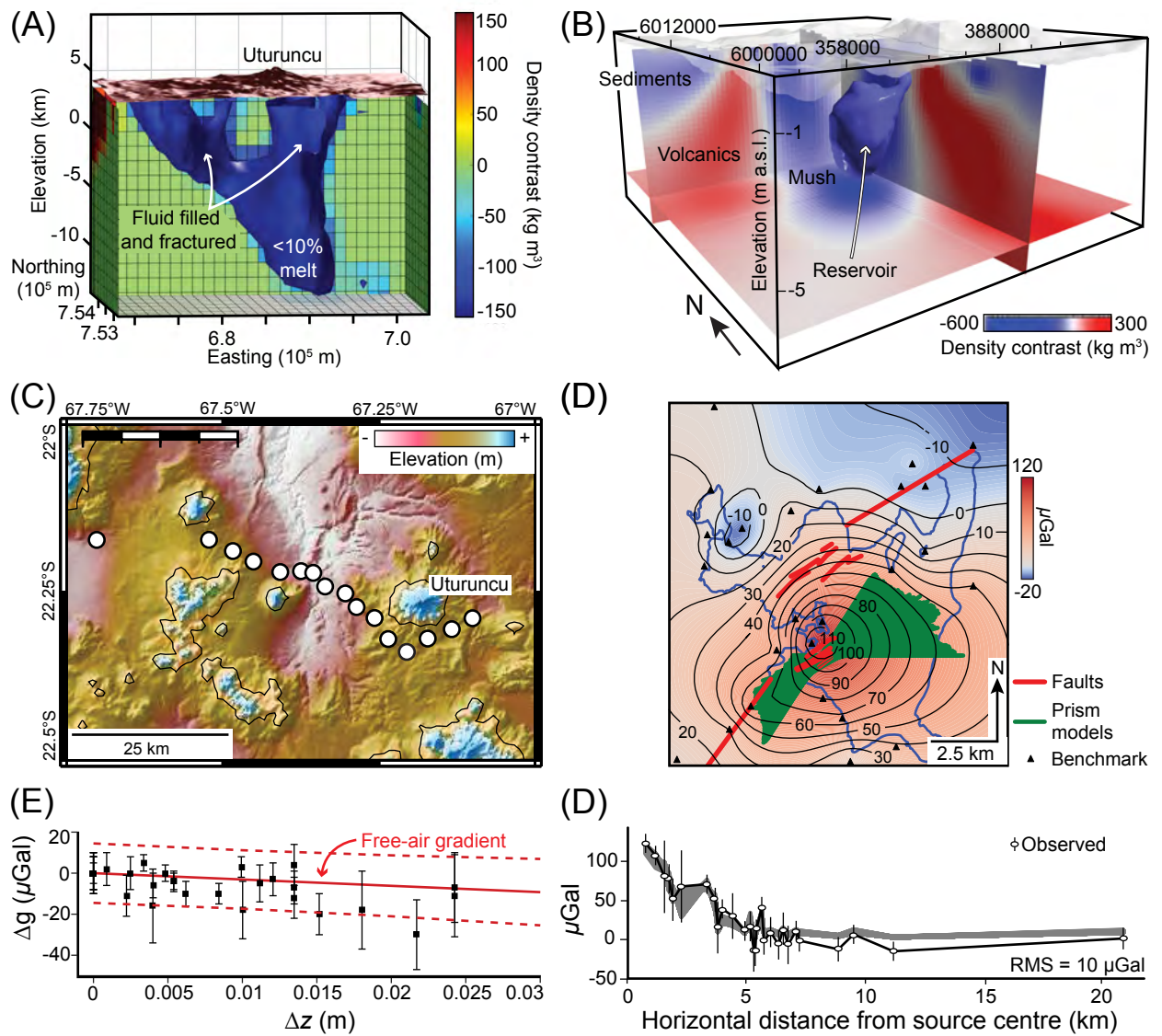


Figure 7



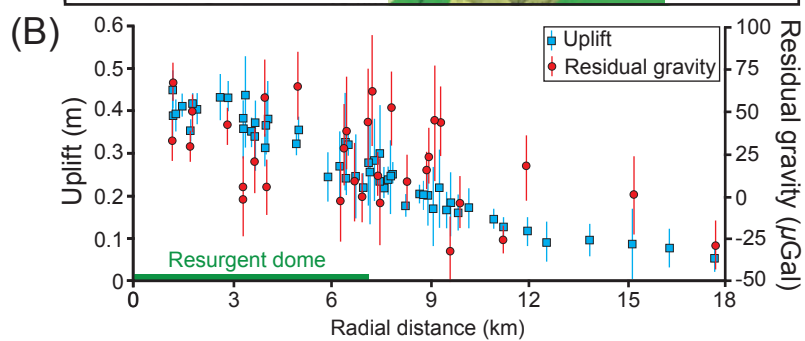
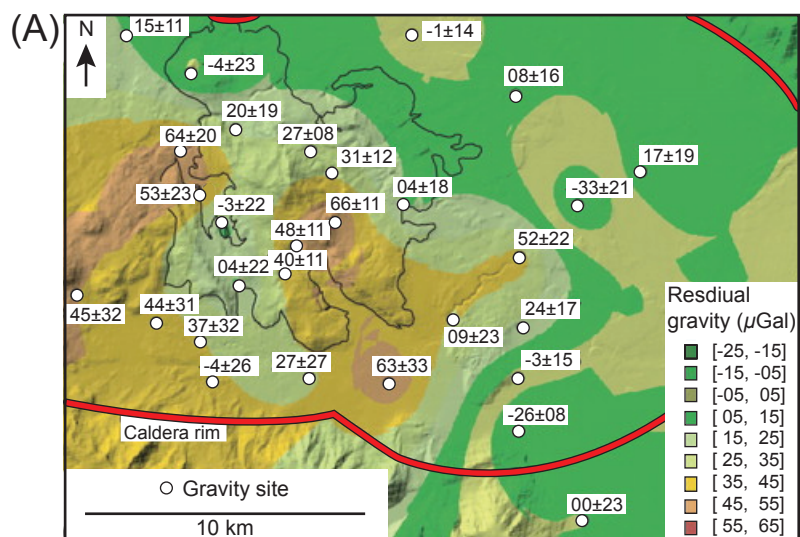


Figure 9

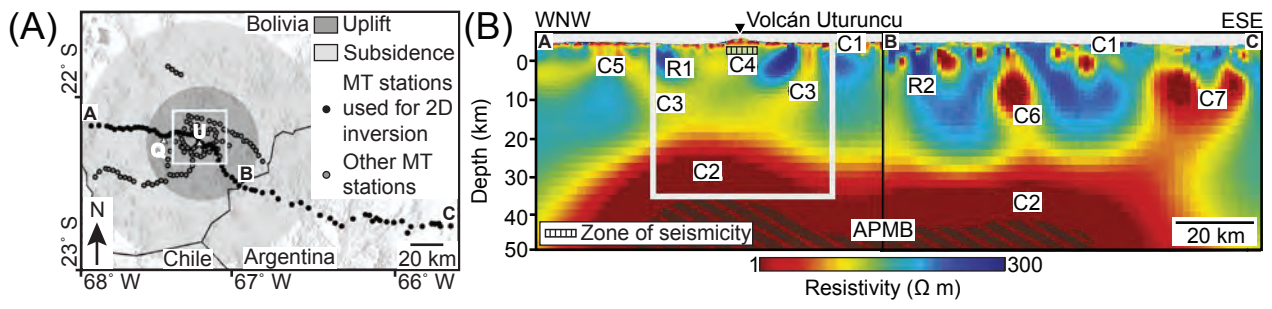




Figure 9

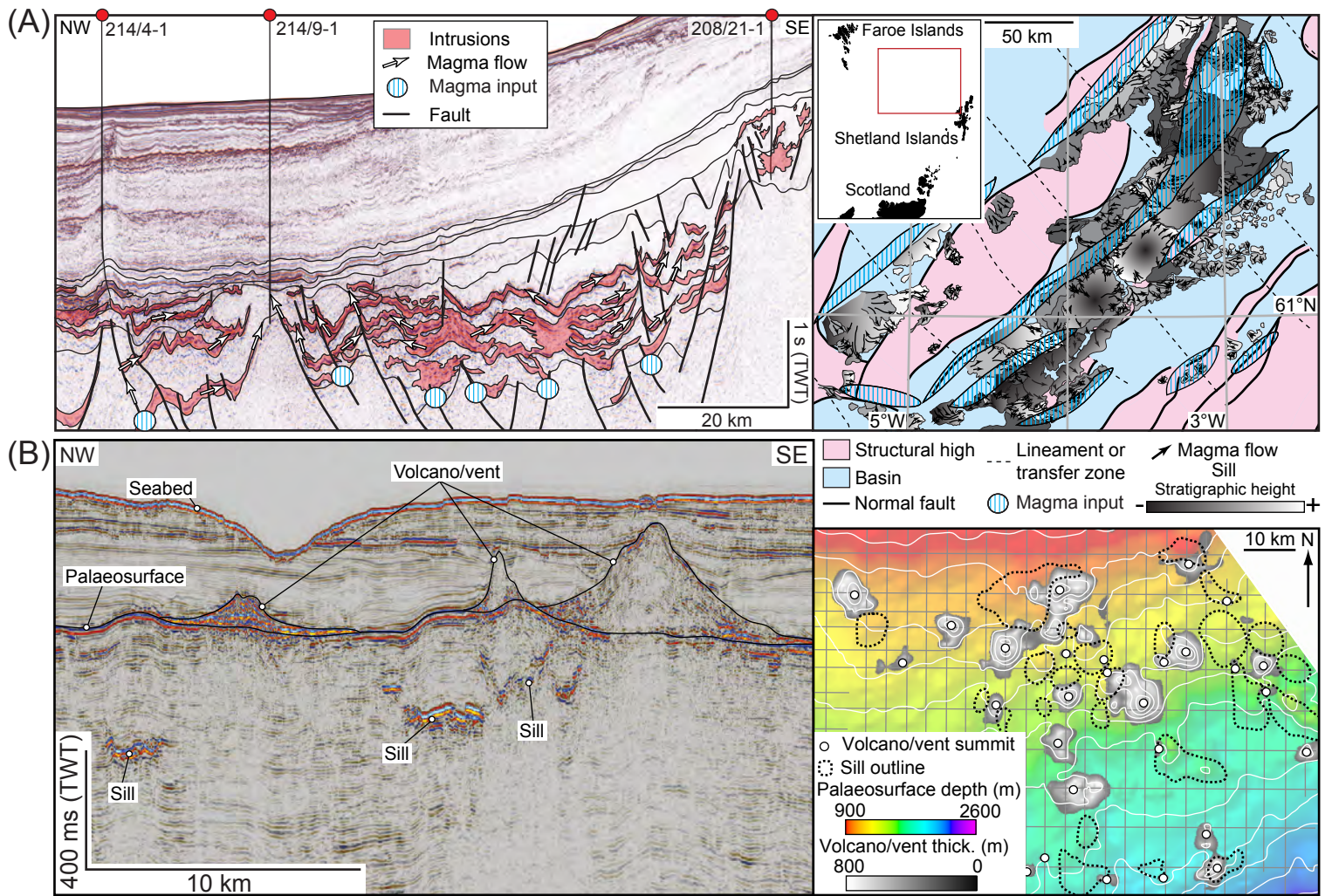


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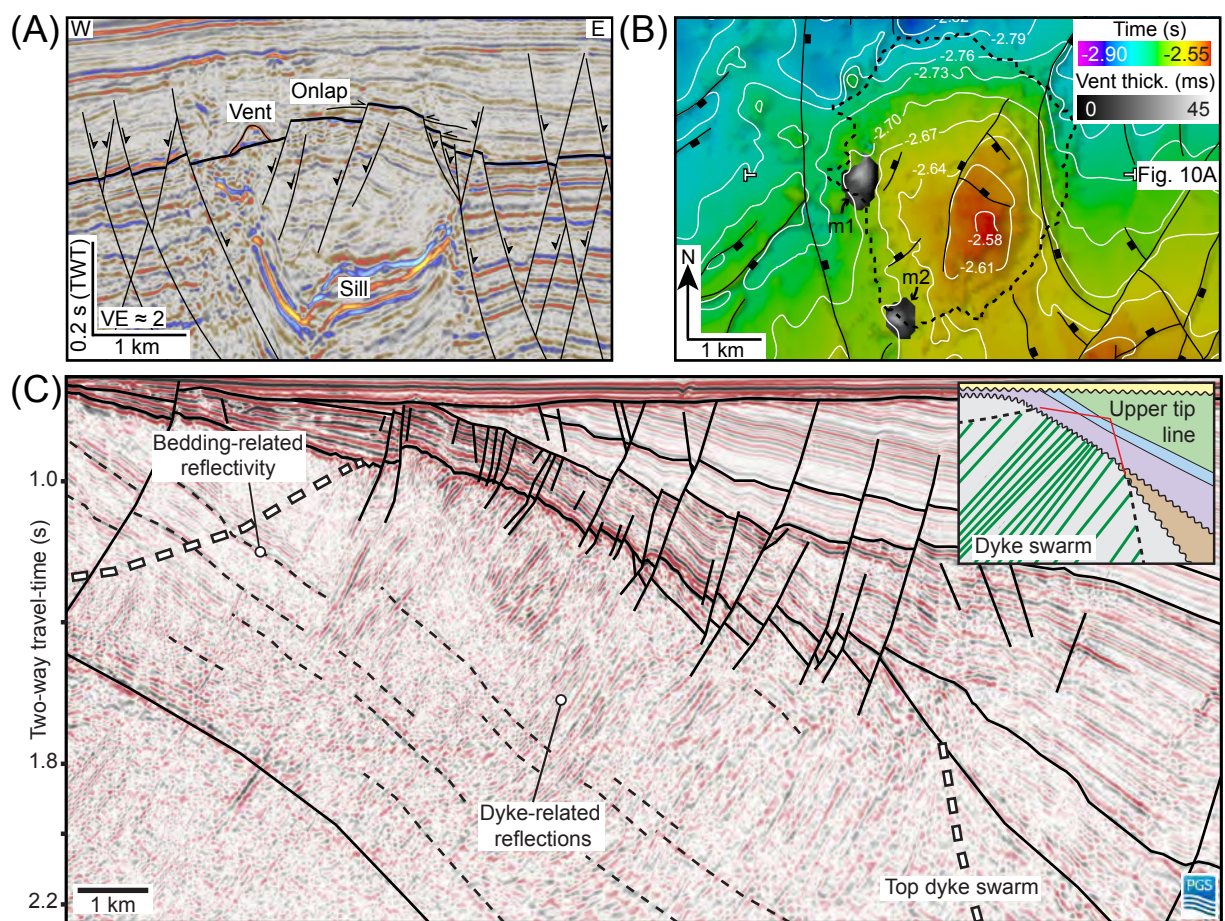




Figure 11

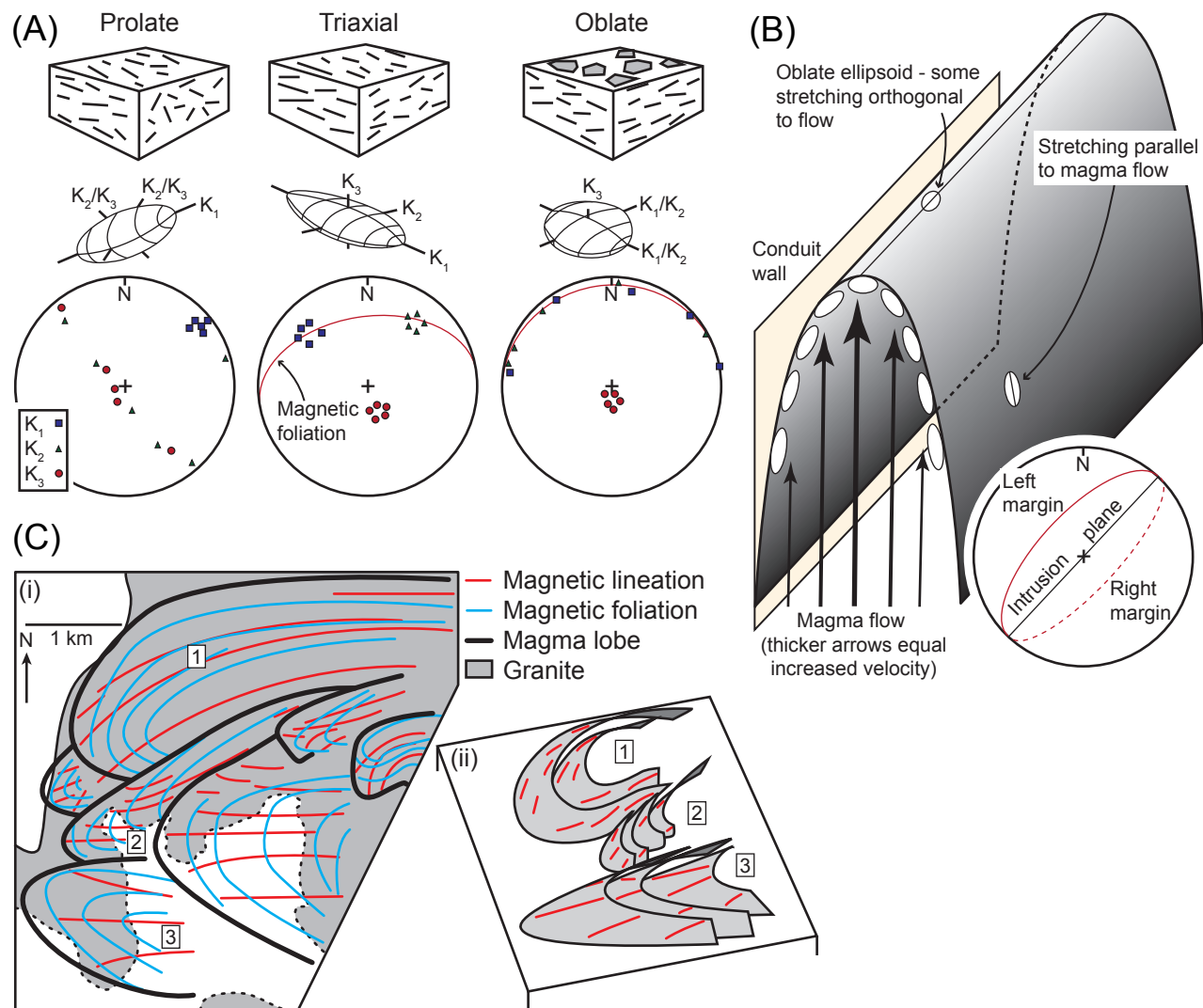




Figure 12

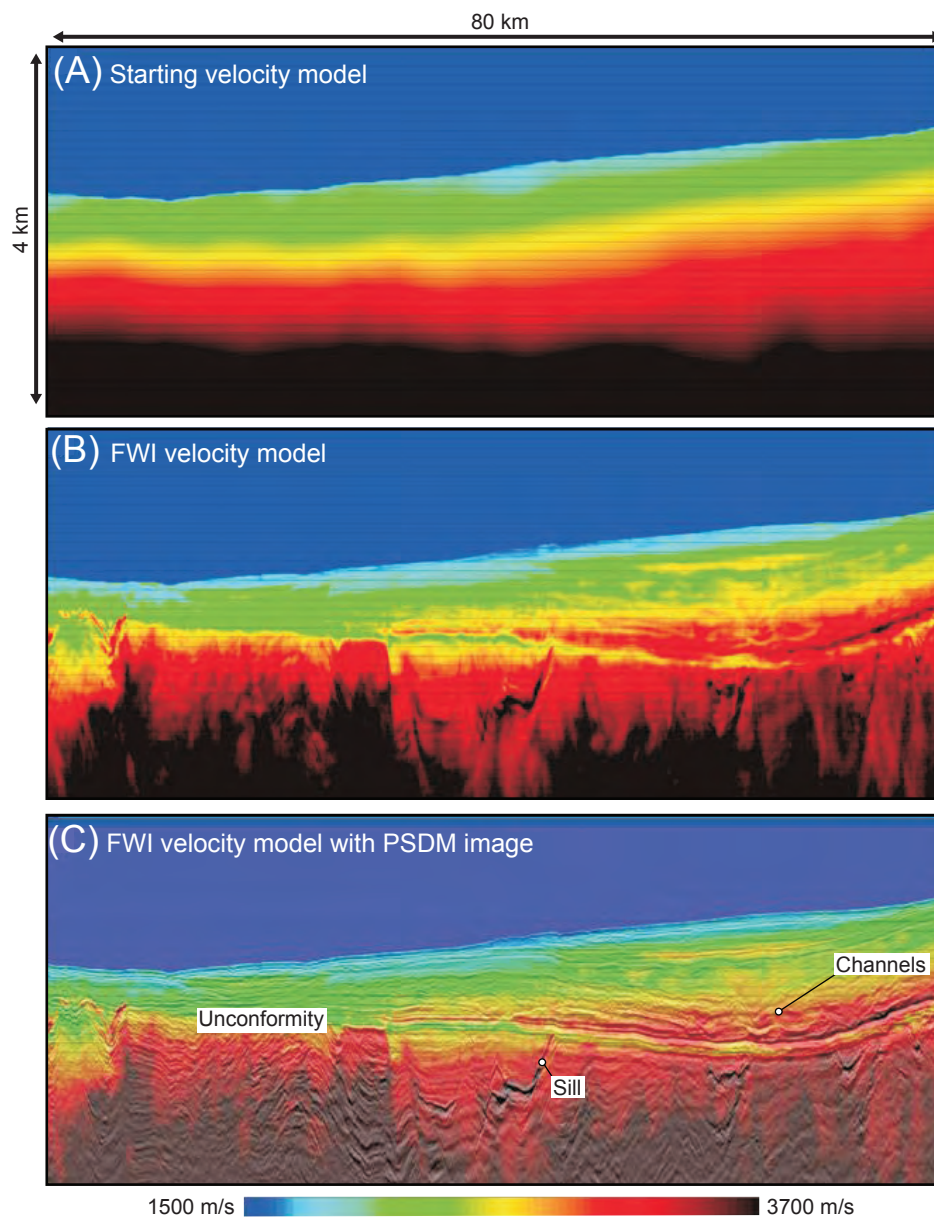


Figure 13

